

VELOCITY CURVE AND MAP OF THE MILKY WAY EPFL



GALACTIC MAP

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INTRODUCTION

The velocity curve of a galaxy displays the velocity at which an object moves at a given distance from the center, assuming it follows a circular trajectory. Such curves require radio astronomical observation using the 21 cm emission line of hydrogen. Velocity curves further served as a main argument in favor of the existence of what is called "Dark Matter" [2]. In this work, such a curve was constructed to investigate the relation with different models for the mass distribution of the Milky Way.

WORKING PRINCIPLE OF THE ANTENNA AND METHOD



Figure 1: VEGA antenna

H21 emission line:

- Spin flip of the electron: rare event
- BUT large abundance of hydrogen in Universe

VEGA (Very Elegant Galactic Antenna):

- Receiver: copper helical antenna around nonconducting rod inside aluminum pan
- Feed: electronic components (amplifiers, filters, bias tee)
- Recorder : electrical → digital converter

Signal processing:

- Obtain a Power Spectral Density curve
- First uniform filtering of the PSD signal
- Substracting the noise

CALIBRATION

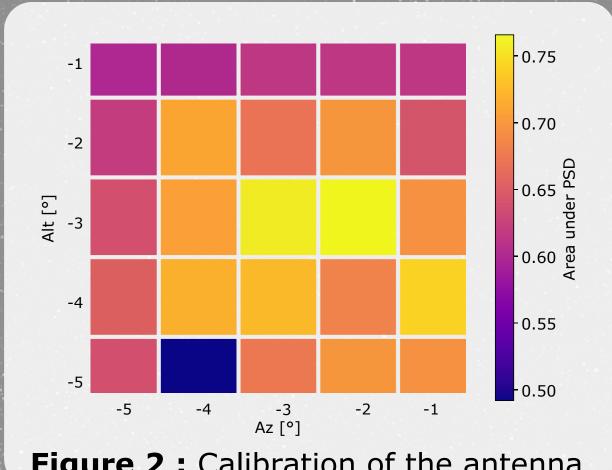


Figure 2: Calibration of the antenna

- Pointing antenna to the exact coordinates of the Sun
- Calculating the area under the PSD curve
- Repeating the operation with an offset on the Az/Alt coordinates to find maximum received power
- As seen on the figure, best offset is $(-2\degree, -3\degree)$ Az/Alt

Velocity-Vector Method:

• From cosine law:

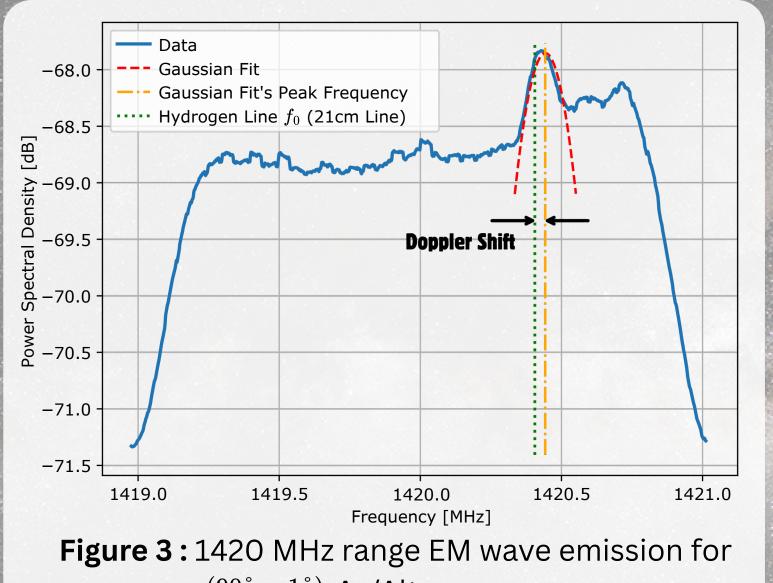
distance:

Approximation for Sun-Cloud

 $R=\sqrt{R_0^2+d^2-2R_0d\cos(l)}$

21 CM SIGNAL

- Hydrogen Emission Line at $f_0 = 1420.405751 \, \mathrm{MHz}$
- Movement of the emitting source (hydrogen cloud)⇒ Doppler shift observed
- Radial velocity (along
- line of sight): $v_r = c \frac{f_{obs} f_0}{f_{obs}}$ (Doppler formula)



 $(90\degree,-1\degree)$ Az/Alt

ullet Formation of a clearer peak for $b o 0^\circ$ (more emission in the milky way than in outer space)

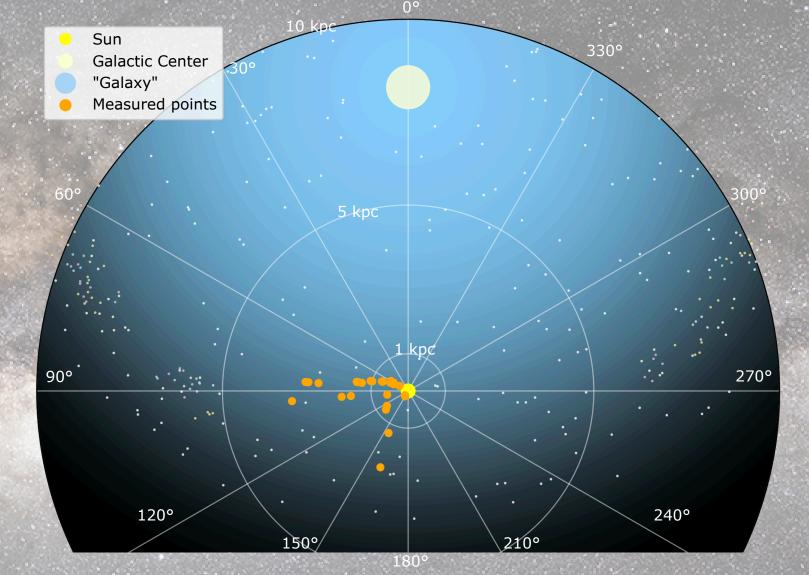


Figure 4: Schematic map of the Galaxy with points measured in this work (local Galactic disk: distance to the Sun <5kpc)

- \rightarrow Oort constant: $A = (14 \pm 1) \text{ km} \cdot \text{s}^{-1} \cdot \text{kpc}^{-1}$
- ightarrow Galactic longitude : $l \in [60^\circ, 160^\circ]$ and latitudes $b = \{-1, 0, 1\}$
- ightarrow Distance Sun Galactic Center: $R_0 pprox 8.17~\mathrm{kpc}$

VELOCITY CURVE

Experimental velocity:

From velocity-vector method :

$$egin{align} v_{circ}\cos\left(igtrule(v_r,v_{circ})
ight) &= v_r + v_0\sin(l) \ \Rightarrow v_{circ} &= rac{R}{R_0}\Big(rac{v_r}{sin(l)} + v_0\Big) \ \end{aligned}$$

ightarrow Sun's velocity around GC : $v_0 pprox 220~{
m km.s}^{-1}$ [4]

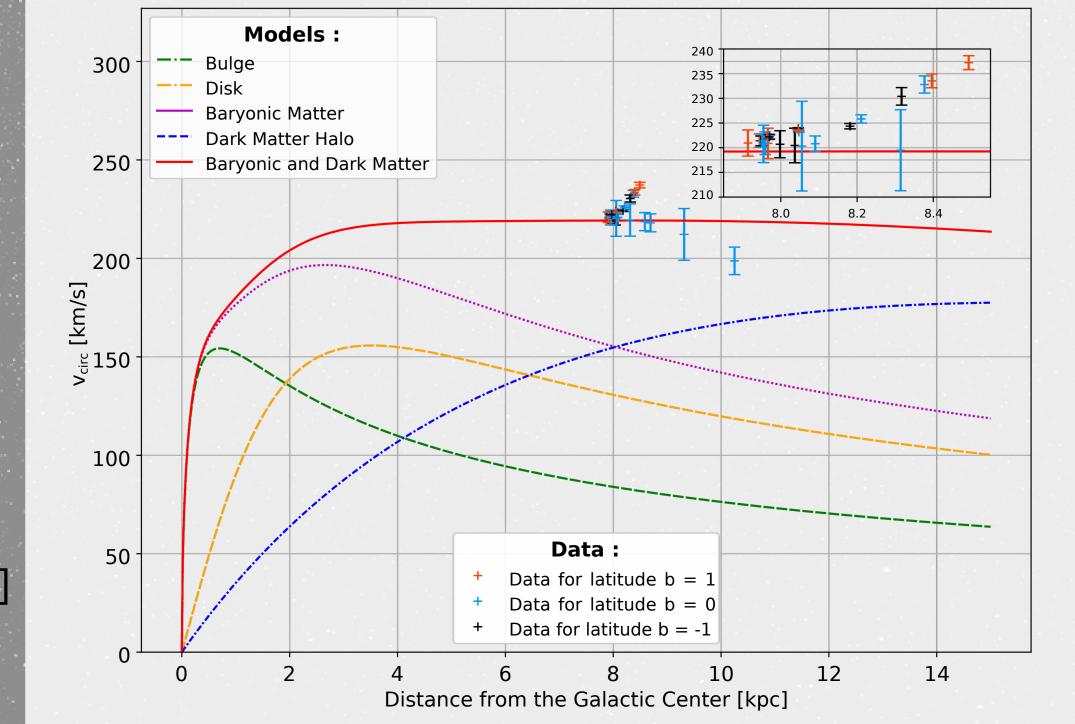
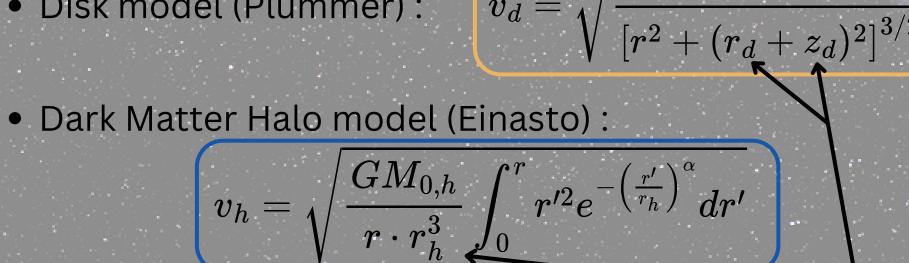


Figure 5: Velocity curve for measured points and theoretical models

Theoretical velocity profiles [5]:

- Circular motion \Rightarrow acceleration : a =gives velocity: $v=\sqrt{-rrac{\partial\phi}{\partial r}}$ Normalized masses ullet Bulge model (Hernquist) : $ig|v_b=\sqrt{r}$
- Disk model (Plummer):



length scales

• Total theoretical velocity curve:

CONCLUSION

In this work, the 21 cm emission line from neutral hydrogen was observed, using the VEGA antenna. This HI line was shown to be a powerful tool for tracing the structure and dynamics of galaxies, including the Milky Way. The former was roughly mapped in the Galactic Map, while the latter was studied through a velocity curve.

These results show the potential of radio observations, even with rather modest instrumentation, to explore large-scale Galactic properties. Further observations at more longitudes would allow for a more complete velocity curve since one of the problems found in this work is the lack of measurements further away from the Galactic Center.

REFERENCES