

Lecture # 10.

Main points of #9.

Considered idea of nucleosynthesis

Freezing temperature of reactions which transforms $p \leftrightarrow n$ is $T^* \approx 0.8 \text{ MeV} \Rightarrow$

$$\frac{n_n}{n_p} = e^{-\frac{m_n - m_p}{T^*}} \approx \frac{1}{8}$$

Then this is somewhat decreased because of n -decays, down to $n_n/n_p \approx \frac{1}{7}$ at $T \approx 0.1 \text{ MeV}$

All n go to ${}^4\text{He}$, as it has largest binding energy for light elements

$$\text{so } X_4 = \frac{4 \cdot n_{\text{He}}}{n_p + n_n} = 0.25$$

Today: Baryogenesis

- observed asymmetry
- candidates for new matter particles
 - $\rightarrow \chi_0, \chi_{\pm}, \tilde{\chi}_{\pm}$
 - $\rightarrow$ gluino, squarks
 - $\rightarrow$ stop squarks
 - $\rightarrow$ gluinos

Baryogenesis

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Questions: why $n_B - n_{\bar{B}} \neq 0$?

why $\frac{n_B - n_{\bar{B}}}{n_\gamma} \approx 10^{-10}$?

- What is needed (Sakharov)

(i) baryon number nonconservation

But' consistent with stability
of matter :

$$\tau_p \approx 10^{31} \text{ years}$$

(ii) Deviations from thermal
equilibrium :

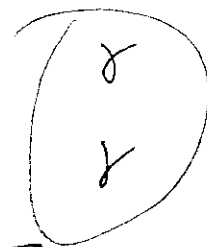
In thermal equilibrium

$$n_p = n_{\bar{p}} = \frac{1}{e^{E/T} - 1}$$

(NO chemical
potential
as baryon
number is
non-conserved
parities

(iii) violation of C & CP

particles : $p \quad n \quad e \quad \nu$
 $\bar{p} \quad \bar{n} \quad e^+ \quad \bar{\nu}$



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truly neutral particle

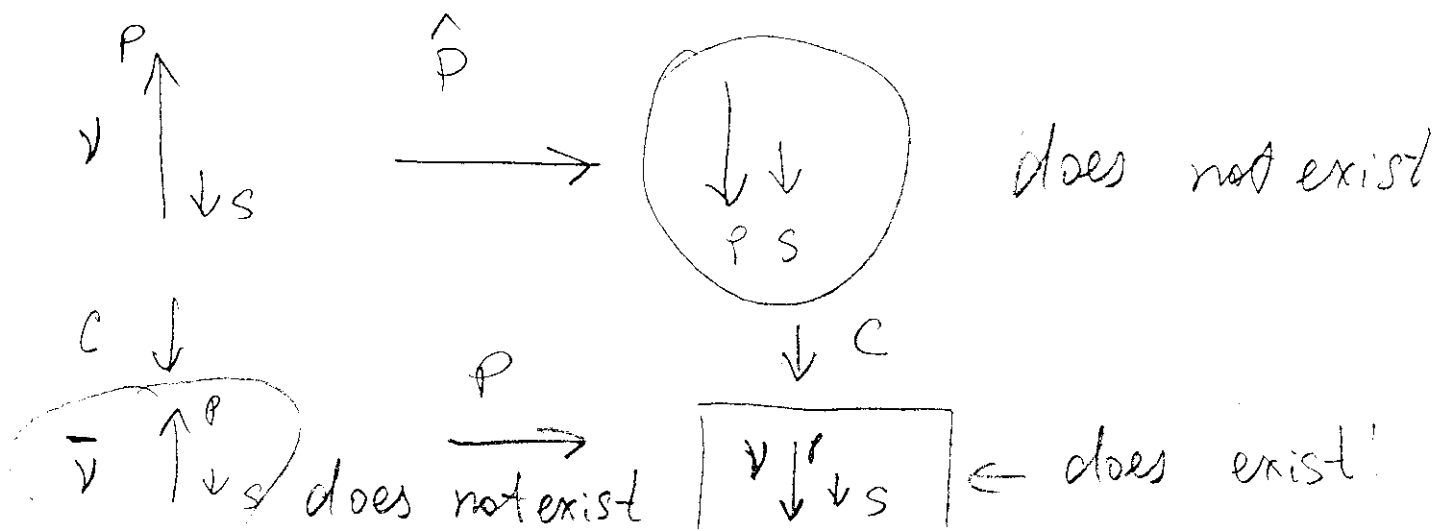
CPT theorem : masses and total lifetimes are equal.

(one of the reasons why in thermal equilibrium baryon number is zero)

Before 1956 : people believed that P & C parities were exact.

$$\left\{ \begin{array}{l} P: x \rightarrow -x, t \rightarrow +t, \quad \underbrace{v \rightarrow -v}_{\text{velocity}} \\ \text{spin} \rightarrow \text{spin} \quad (s \sim \vec{x} \times \vec{v}) \\ C: \text{particle to antiparticle} \end{array} \right.$$

Discovery of P & C non-conservation associated with neutrinos



So, P and C are broken, but CP may be exact

If CP is exact - no difference between particles and antiparticles $\Rightarrow$ baryon asymmetry cannot arise

1964 - discovery of CP -breaking in decays of K^0 mesons

(K_1 & K_2 , K_S & K_L)

$$\begin{array}{ll}
 K^0 \rightarrow \pi^- e^+ \nu & \text{asymmetry in } K^0 \text{ decays} \\
 \bar{K}^0 \rightarrow \pi^+ e^- \bar{\nu} & \frac{\# e^+ - \# e^-}{\# e^+ + \# e^-} \sim O(10^{-3})
 \end{array}$$

"Absolute definition" of matter

- take equal number of K^0 & $\bar{K}^0$
- call "positron" a charged lepton which appears more often in decays
- call proton the baryon of the same sign of charge



If CP is broken then baryon asymmetry can arise.

Grand Unified baryogenesis

Grand unified theories — (GUT)
unification of strong, weak & electromagnetic interactions

Information on GUT's important for us (for explanation of Baryogenesis)

fundamental particles — baryons — are quarks rather than protons or neutrons

$$q = \begin{pmatrix} u \\ d \end{pmatrix} \quad \begin{pmatrix} c \\ s \end{pmatrix} \quad \begin{pmatrix} t \\ b \end{pmatrix}$$

$$q = \frac{2}{3}, \quad b = \frac{1}{3}$$

$$q = -\frac{1}{3}, \quad b = \frac{1}{3}$$

electric charge baryon number

nucleons consist of 3 quarks:

$$p = uud$$

$$n = udd$$

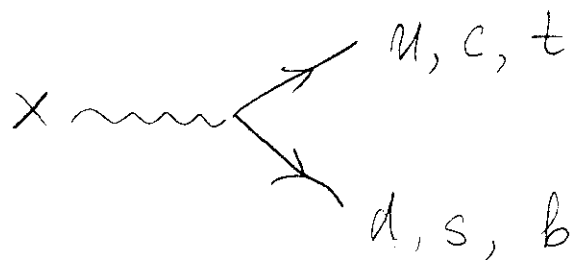
mesons are constructed from quarks & antiquarks, $q\bar{q}$

(small temperatures & densities)

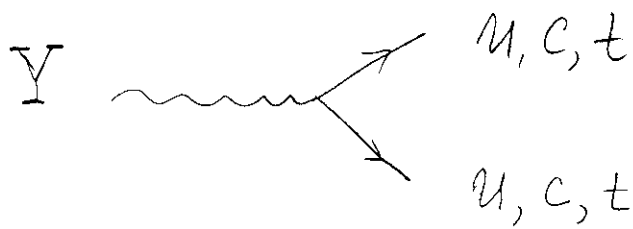
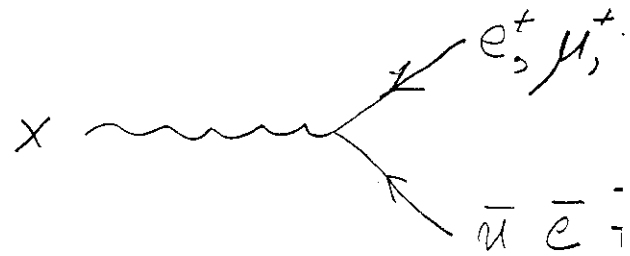
In ordinary conditions quarks do not exist in free state (confinement of quarks)

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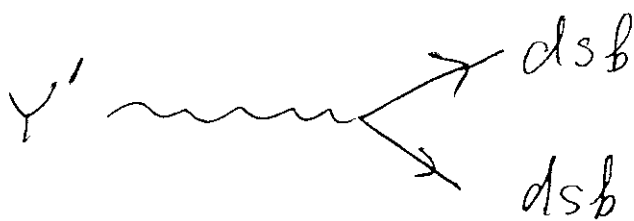
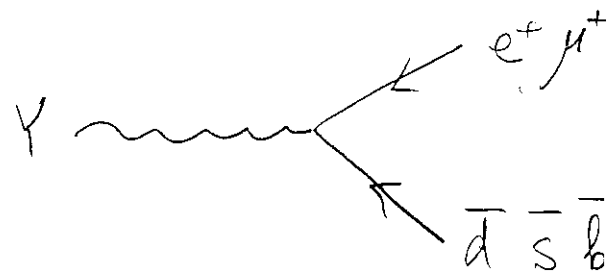
In GUTs there are particles -
leptoquarks - which can decay on
quarks and leptons with baryon
number non-conservation:



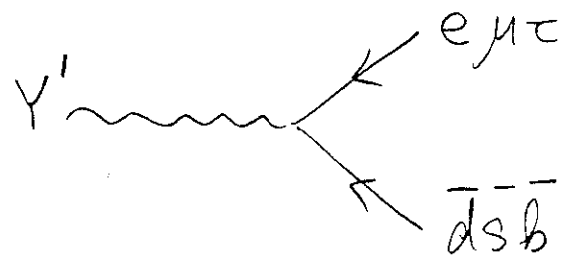
$$Q_X = \frac{1}{3}$$



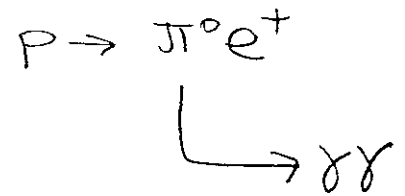
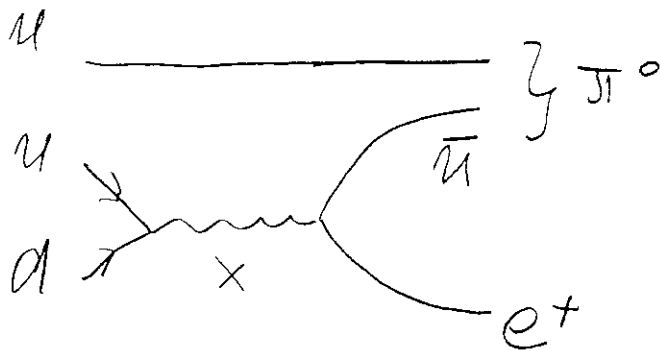
$$Q_Y = \frac{4}{3}$$



$$Q_{Y'} = -\frac{2}{3}$$



proton is unstable:



proton width:

$$\Gamma = \frac{g^4}{M_X^4} \cdot m_p^5 \approx \frac{\alpha^2 m_p^5}{M_X^4}$$

Home work: find constraint on M_X from $\tau_p \geq 10^{31}$ years.

CP - violation in leptoquark decays

$$X \text{ (wavy line)} \rightarrow \begin{matrix} q \\ q \end{matrix} : \Gamma(X \rightarrow qq)$$

$$X \text{ (wavy line)} \rightarrow \begin{matrix} \bar{q} \\ \bar{e} \end{matrix} : \Gamma(X \rightarrow \bar{q} \bar{e})$$

$$\bar{X} \text{ (wavy line)} \rightarrow \begin{matrix} \bar{q} \\ q \end{matrix} : \Gamma(\bar{X} \rightarrow \bar{q} q)$$

$$\bar{X} \text{ (wavy line)} \rightarrow \begin{matrix} q \\ e \end{matrix} : \Gamma(\bar{X} \rightarrow q e)$$

CPT :

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$$\Gamma(X \rightarrow qq) + \Gamma(X \rightarrow \bar{q}\bar{\ell}) = \Gamma_{tot} = \\ = \Gamma(\bar{X} \rightarrow \bar{q}\bar{q}) + \Gamma(\bar{X} \rightarrow q\ell)$$

$$\text{or: } \Gamma(X \rightarrow qq) - \Gamma(\bar{X} \rightarrow \bar{q}\bar{q}) = \Gamma(\bar{X} \rightarrow q\ell) - \Gamma(X \rightarrow \bar{q}\bar{\ell})$$

CP-violation :

$$\Gamma(X \rightarrow qq) \neq \Gamma(\bar{X} \rightarrow \bar{q}\bar{q})$$

Suppose we have one X and one $\bar{X}$

How many baryons we will get after their decay?

$$X : B = \frac{2}{3} \frac{\Gamma(X \rightarrow qq)}{\Gamma_{tot}} - \frac{1}{3} \frac{\Gamma(X \rightarrow \bar{q}\bar{\ell})}{\Gamma_{tot}}$$

$$\bar{X} : B = -\frac{2}{3} \frac{\Gamma(\bar{X} \rightarrow \bar{q}\bar{q})}{\Gamma_{tot}} + \frac{1}{3} \frac{\Gamma(\bar{X} \rightarrow q\ell)}{\Gamma_{tot}}$$

$$\text{total } B : \frac{2}{3} \frac{\Gamma(X \rightarrow qq) - \Gamma(\bar{X} \rightarrow \bar{q}\bar{q})}{\Gamma_{tot}} - \frac{1}{3} \frac{\Gamma(X \rightarrow \bar{q}\bar{\ell}) - \Gamma(\bar{X} \rightarrow q\ell)}{\Gamma_{tot}}$$

$$= \frac{\Gamma(X \rightarrow qq) - \Gamma(\bar{X} \rightarrow \bar{q}\bar{q})}{\Gamma_{tot}}$$

Baryon asymmetry

$$\Delta = \frac{n_B}{S} \sim \frac{\Gamma(x \rightarrow qq) - \Gamma(\bar{x} \rightarrow \bar{q}\bar{q})}{\Gamma_{tot}} \cdot \frac{n_x}{S}$$

concentration $\propto$ leptogluarks 9

$\nearrow$
entropy density

Kinetics & thermodynamics of
leptogluark decay

(or, how to find n_x/S).

Thermal equilibrium:

$$\frac{n_x}{n_f} \approx \begin{cases} \frac{T^3}{T^3} \approx 1 & , \text{ if } T \gg m_x \\ \left(\frac{m}{T}\right)^{3/2} \exp\left(-\frac{m_x}{T}\right) & , \text{ if } T \ll m_x \end{cases}$$

processes in which leptogluarks take part:

$$X \rightarrow qq ; X \rightarrow \bar{q}\bar{e}$$

and (very important)

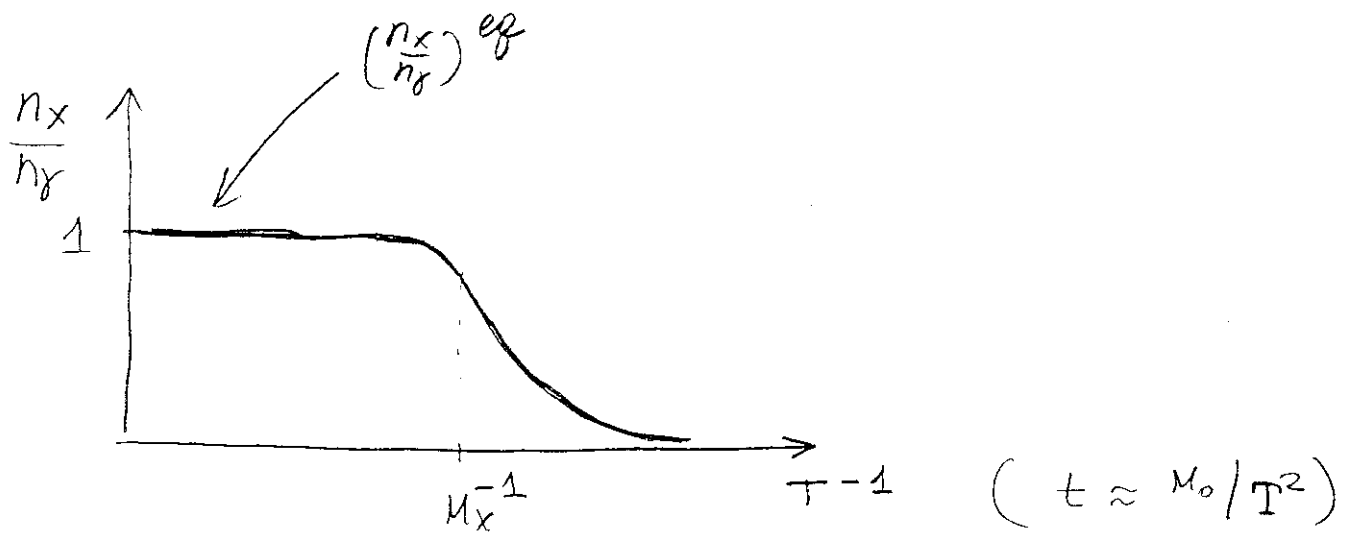
inverse processes

$$qq \rightarrow X ; \bar{q}\bar{e} \rightarrow X$$

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in thermal equilibrium rates of direct and inverse processes are the same
rate of direct processes:

$$\Gamma_{\text{tot}} \approx \begin{cases} \Gamma_{\text{tot}} \frac{m_x}{T} & , \text{ if } T \gg m_x \\ \Gamma_{\text{tot}} & \text{ if } T \lesssim m_x \end{cases} \quad \frac{T}{m_x} - \text{Lorentz factor}$$



decoupling temperature:

$$(*) \quad \Gamma_{\text{tot}} \approx \frac{T^2}{M_0} \quad ; \quad T^* \approx \sqrt{\Gamma_{\text{tot}} M_0} \quad (\text{if } T^* \lesssim m_x)$$

$$(**) \text{ or } \Gamma_{\text{tot}} \frac{m_x}{T} \approx \frac{T^2}{M_0} \quad , \quad T^* \approx (\Gamma_{\text{tot}} m_x M_0)^{1/3} \quad (\text{if } T^* \gtrsim m_x)$$

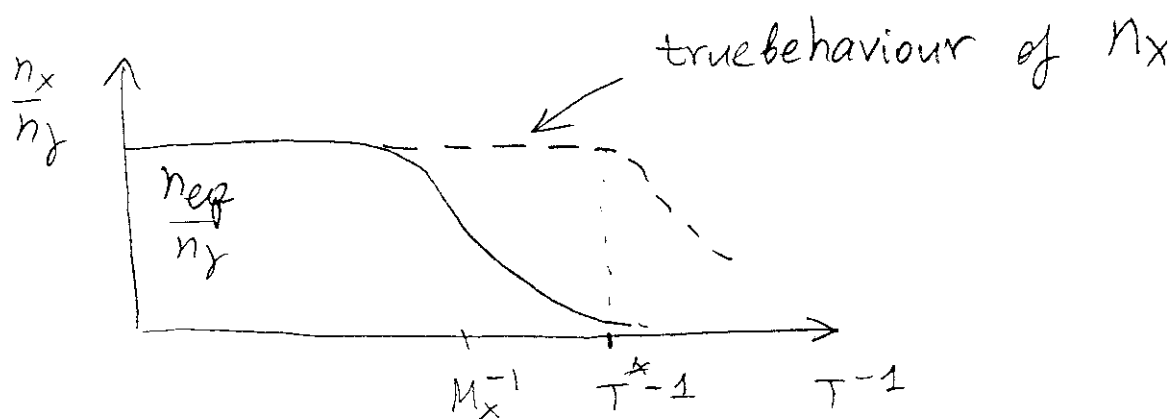
$$(*) \quad \text{1st case: } \Gamma_{\text{tot}} M_0 \lesssim m_x^2 \quad ; \quad \frac{m_x^2}{\Gamma_{\text{tot}} M_0} \gtrsim 1$$

$$(**) \quad \text{2nd case: } \frac{m_x^2}{\Gamma_{\text{tot}} M_0} \lesssim 1$$

$$\boxed{\text{let } T^* \lesssim M_X} \Rightarrow$$

$$\frac{n_X}{n_\gamma} = 1 \quad \text{till } T^*, \text{ because}$$

particles do not decay at all at $T \gtrsim T^*$



in this case $\frac{n_X}{n_\gamma} \approx 1$ at the moment of decay $\Rightarrow$

$$\Delta = \frac{n_B}{S} \approx \frac{\Gamma(X \rightarrow qq) - \Gamma(\bar{X} \rightarrow \bar{q}\bar{q})}{\Gamma_{tot}} \cdot \frac{1}{N_{eff}}$$

N_{eff} - number of degrees of freedom,

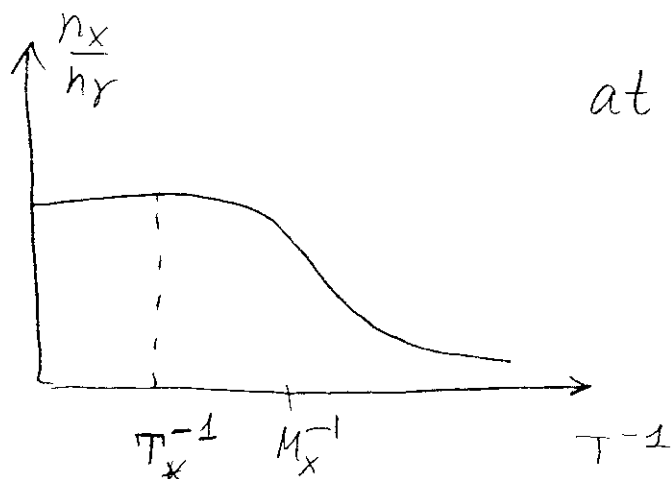
$$N_{eff} \sim 100$$

(counting all quarks, leptons, vector boson)

$$\text{Denote } \delta_{micro} \equiv \frac{\Gamma(X \rightarrow qq) - \Gamma(X \rightarrow \bar{q}\bar{q})}{\Gamma_{tot}} \frac{1}{N_{eff}}$$

Let $\frac{*}{T} \approx M_X$

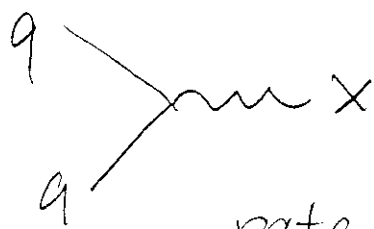
$$\frac{n_X}{n_\gamma} \approx \frac{n_{eq}}{n_\gamma} \text{ till } \frac{*}{T}$$



at $T \leq T^*$ rate of decay
is larger than the rate
of Universe expansion

Computation is more complicated

Consider inverse process



rate of inverse process :

$$\Gamma_{\text{invers}} \sim \Gamma_{\text{tot}} \underbrace{\exp\left(-\frac{m_X}{T}\right) \left(\frac{m_X}{T}\right)^{3/2}}_{\text{factor}}$$

→ this factor comes from concentration
of quarks, $n_q(1) \cdot n_q(2) \sim \exp\left(-\frac{E_{\text{tot}}}{T}\right)$,
 $E_{\text{tot}} = m_X$ in order to create X

Compare it with the rate of universe expansion:

$$\left(\frac{m_x}{T}\right)^{3/2} \Gamma_{tot} \exp\left(-\frac{m_x}{T}\right) \approx \frac{T^2}{M_0} \Rightarrow$$

$$\left(\frac{m_x}{T}\right) = x; \quad x^{7/2} e^{-x} = \frac{m_x^2}{M_0 \Gamma_{tot}}$$

$$\frac{7}{2} \log x - x = \log \frac{m_x^2}{M_0 \Gamma_{tot}};$$

$$x \approx \log \frac{M_0 \Gamma_{tot}}{m_x^2} + \frac{7}{2} \log \log \frac{M_0 \Gamma_{tot}}{m_x^2} \Rightarrow$$

$$\frac{n_x}{n_\gamma} \approx x^{3/2} e^{-x} \approx$$

$$\approx \left[\log \frac{M_0 \Gamma_{tot}}{m_x^2} \right]^{3/2} \frac{m_x^2}{M_0 \Gamma_{tot}} \left[\log \frac{M_0 \Gamma_{tot}}{m_x^2} \right]^{-7/2} =$$

$$= \frac{m_x^2}{M_0 \Gamma_{tot}} \frac{1}{\left(\log \frac{M_0 \Gamma_{tot}}{m_x^2} \right)^2}$$

Note: Solution to Boltzmann equation gives log in the 1st power.

So :

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$$\Delta \approx \begin{cases} \delta_{\text{micro}} & ; \quad \frac{m_x^2}{\Gamma_{\text{tot}} M_0} \gtrsim 1 \\ \delta_{\text{micro}} \cdot \frac{m_x^2}{\Gamma_{\text{tot}} M_0} & ; \quad \frac{m_x^2}{\Gamma_{\text{tot}} M_0} \lesssim 1 \end{cases}$$

Usually : $m_x \approx \alpha \cdot M_0$, $\alpha \approx \alpha_{\text{EM}} \approx O\left(\frac{1}{100}\right)$

δ_{micro} : at most $O\left(\frac{1}{N_{\text{eff}}}\right) \sim \frac{1}{100}$
if CP-violation is 100%

$$\Rightarrow \frac{m_x^2}{\Gamma_{\text{tot}} M_0} \gtrsim 10^{-8} ;$$

$$m_x \gtrsim 10^{-8} \alpha M_0 \approx 10^8 \text{ GeV}$$

if CP-violation is small,

$\delta_{\text{micro}} \sim 10^{-10}$, then one
requires $m_x \gtrsim \alpha M_0 \approx 10^{16} \text{ GeV}$