

General Relativity Formulary

1) Differential geometry

Metric connection (Christoffel symbols)

$$\Gamma_{\beta\gamma}^\alpha \equiv \frac{1}{2} g^{\alpha\delta} (\partial_\beta g_{\gamma\delta} + \partial_\gamma g_{\beta\delta} - \partial_\delta g_{\beta\gamma}).$$

Covariant derivative

$$V^\alpha_{;\beta} = V^\alpha_{,\beta} + \Gamma_{\beta\gamma}^\alpha V^\gamma$$

$$V_{\alpha;\beta} = V_{\alpha,\beta} - \Gamma_{\alpha\beta}^\gamma V_\gamma.$$

Riemann curvature tensor

$$R^\alpha_{\beta\gamma\delta} \equiv \partial_\gamma \Gamma_{\beta\delta}^\alpha - \partial_\delta \Gamma_{\beta\gamma}^\alpha + \Gamma_{\beta\delta}^\lambda \Gamma_{\lambda\gamma}^\alpha - \Gamma_{\beta\gamma}^\lambda \Gamma_{\delta\lambda}^\alpha,$$

or, with the first index lowered

$$R_{\alpha\beta\gamma\delta} = g_{\alpha\lambda} R^\lambda_{\beta\gamma\delta}$$

$$= \frac{1}{2} (\partial_\beta^2 g_{\alpha\delta} + \partial_\alpha^2 g_{\beta\gamma} - \partial_\alpha^2 g_{\beta\delta} - \partial_\beta^2 g_{\alpha\gamma})$$

$$+ g_{\mu\nu} (\Gamma_{\alpha\delta}^\mu \Gamma_{\beta\gamma}^\nu - \Gamma_{\alpha\gamma}^\mu \Gamma_{\beta\delta}^\nu).$$

Ricci tensor and Ricci scalar

$$R_{\alpha\beta} \equiv R^\gamma_{\alpha\gamma\beta}, \quad R \equiv g^{\alpha\beta} R_{\alpha\beta}.$$

Bianchi identities

$$R_{\alpha\beta\gamma\delta;\lambda} + R_{\alpha\beta\lambda\gamma;\delta} + R_{\alpha\beta\delta\lambda;\gamma} = 0.$$

2) General Relativity

Geodesic equations

$$\frac{d^2 x^\alpha}{dp^2} + \Gamma_{\beta\gamma}^\alpha \frac{dx^\beta}{dp} \frac{dx^\gamma}{dp} = 0.$$

Einstein equations

$$G_{\mu\nu} = 8\pi G T_{\mu\nu},$$

where $G_{\mu\nu}$ is the Einstein tensor defined by

$$G_{\mu\nu} \equiv R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu},$$

and $T_{\mu\nu}$ is the energy-momentum tensor given by

$$T_{\mu\nu} \equiv \frac{2}{\sqrt{-g}} \frac{\delta S_{\text{matter}}}{\delta g^{\mu\nu}}.$$

Variation of the determinant of the metric

$$\delta g = g g^{\mu\nu} \delta g_{\mu\nu} = -g g_{\mu\nu} \delta g^{\mu\nu}.$$

3) Energy-momentum tensor of point-particle

$$T^{\mu\nu} = \frac{m}{\sqrt{-g}} \int d\tau \frac{dz^\mu}{d\tau} \frac{dz^\nu}{d\tau} \delta^4(x - z(\tau)).$$

4) Schwarzschild solution

$$ds^2 = \left(1 - \frac{2GM}{r}\right) dt^2 - \frac{1}{1 - \frac{2GM}{r}} dr^2 - r^2 (d\theta^2 + \sin^2\theta d\phi^2).$$

Geodesic equations for the Schwarzschild geometry

$$\ddot{t} + \frac{B'}{B} \dot{t} \dot{r} = 0$$

$$\ddot{r} + \frac{1}{2} \frac{B'}{A} \dot{t}^2 + \frac{1}{2} \frac{A'}{A} \dot{r}^2 - \frac{r}{A} \dot{\theta}^2 - \frac{r \sin^2\theta}{A} \dot{\phi}^2 = 0$$

$$\ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \sin \theta \cos \theta \dot{\phi}^2 = 0$$

$$\ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot \theta \dot{\theta} \dot{\phi} = 0,$$

where

$$B(r) = 1 - \frac{2GM}{r}, \quad A(r) = B^{-1}(r).$$

Motion in the Schwarzschild metric

$$A(r) \left(\frac{dr}{d\phi} \right)^2 \frac{J^2}{r^4} + \frac{J^2}{r^2} - \frac{1}{B(r)} = -E$$

$$d\tau^2 = EB^2(r)dt^2, \quad r^2 \frac{d\phi}{dt} = JB(r)$$

Massive particle

$$\frac{1}{2} \left(\frac{dr}{d\tau} \right)^2 + \frac{1}{2} \left(1 - \frac{2GM}{r} \right) \left(1 + \frac{J^2/E}{r^2} \right) = \frac{1}{2E}$$

Massless particle ($d\lambda = B(r)dt$)

$$\frac{1}{2} \left(\frac{dr}{d\lambda} \right)^2 + \frac{1}{2} \left(1 - \frac{2GM}{r} \right) \frac{J^2}{r^2} = \frac{1}{2}$$

5) Linearized theory

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$R_{\mu\nu}^{(1)} = \frac{1}{2} (-\square h_{\mu\nu} + \partial_\lambda \partial_\mu h^\lambda_\nu + \partial_\lambda \partial_\nu h^\lambda_\mu - \partial_\mu \partial_\nu h)$$

$$R^{(1)} = -\square h + \partial_\mu \partial_\nu h^{\mu\nu}$$

Einstein equations for physical amplitudes

$$\square h_{12} = 8\pi G T_{12}$$

$$\square h = 4\pi G (T_{22} - T_{11})$$

6) Numerical values

$$G = 6.67 \cdot 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$$

$$c \simeq 3 \cdot 10^8 \text{ m/s}$$

$$M_\odot \simeq 2 \cdot 10^{30} \text{ kg}, \quad r_\odot \simeq 7 \cdot 10^5 \text{ km}$$

$$M_\oplus \simeq 6 \cdot 10^{24} \text{ kg}, \quad r_\oplus \simeq 6000 \text{ km}$$

$$M_{NS} \simeq 2M_\odot, \quad r_{NS} \simeq 12 \text{ km}$$