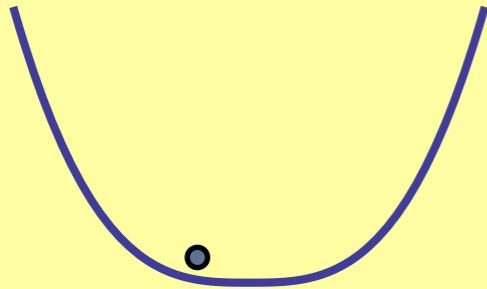


# Laws, Phenomena & History

Riccardo Rattazzi, EPFL

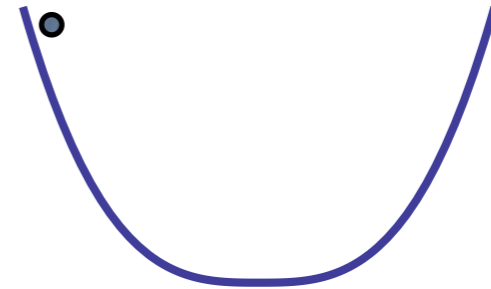
Laws



add quanta



Phenomena



go back in time

add gravity

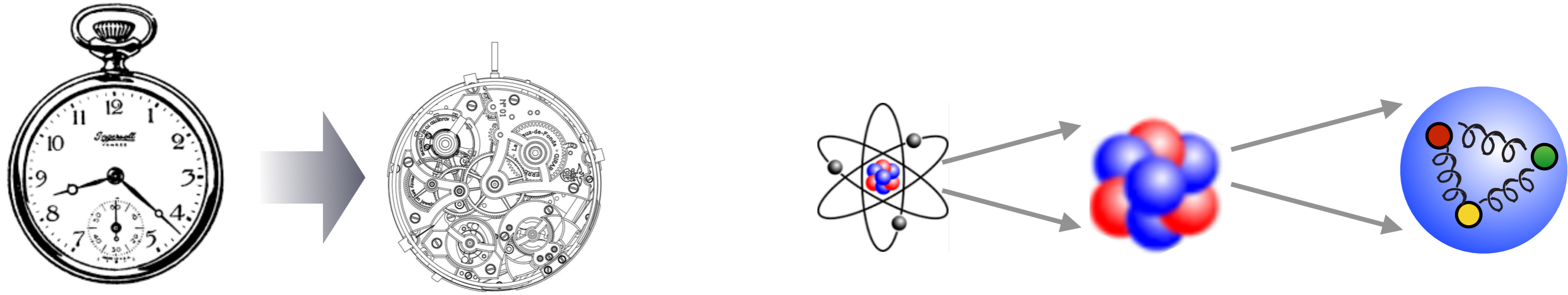
$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = T_{\mu\nu}$$

History

# Particle Physics in a Nutshell

## The Hierarchy Paradox

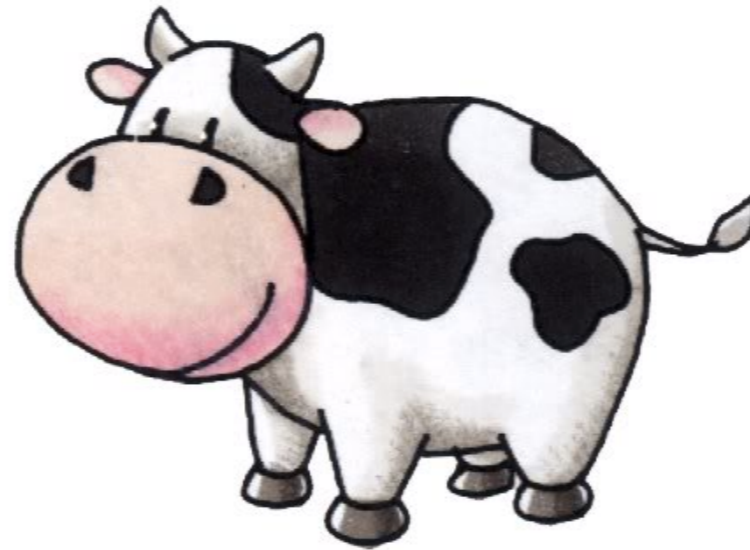
# ▲ Reductionism



# ▲ Effective Long Distance Description

- *Multipole expansion*
- *Effective Field Theory*
- ...

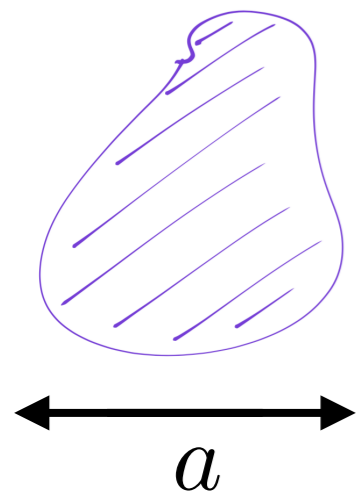
# Long Distance Physics: Simplicity & Accidental Symmetries



accidental

$$SO(3)$$

Ex.: electrostatic potential at large distance



$$\xrightarrow{R \gg a}$$

$$\Phi(R) = \frac{1/R}{R} + \frac{a/R^2}{R^3} + \frac{a^2/R^3}{R^5} + \dots$$

$$SO(3) \supset SO(2) \supset \emptyset$$

Modern view

Standard Model is just an effective field theory  
valid below a physical energy cut-off  $\Lambda_{UV} = 1/a$

( $E \ll \Lambda_{UV}$  or  $\lambda \gg \Lambda_{UV}^{-1}$ )

$$\mathcal{L}_{SM} = \mathcal{L}^{d \leq 4} + \frac{1}{\Lambda_{UV}} \mathcal{L}^{d=5} + \frac{1}{\Lambda_{UV}^2} \mathcal{L}^{d=6} + \dots$$

the three problems

$$+ \Lambda_{UV}^4 \sqrt{g}$$

d=0

$$+ c\Lambda_{UV}^2 H^\dagger H$$

d=2

$$+ \theta \tilde{G}_{\mu\nu} \tilde{G}^{\mu\nu}$$

d=4

$$\mathcal{L}_{SM} = \mathcal{L}_{kin} + gA_\mu \bar{F} \gamma_\mu F + Y_{ij} \bar{F}_i H F_j + \lambda(H^\dagger H)^2$$

d=4

$$+ \frac{b_{ij}}{\Lambda_{UV}} L_i L_j H H$$

$$+ \frac{c_{ijkl}}{\Lambda_{UV}^2} \bar{F}_i F_j \bar{F}_k F_\ell + \frac{c_{ij}}{\Lambda_{UV}} \bar{F}_i \sigma_{\mu\nu} F_j G^{\mu\nu} + \dots$$

$$+ \dots$$

d>4

$\Lambda_{UV} \gg \text{TeV}$  (pointlike limit) nicely accounts for 'what we see'

# The Hierarchy Paradox

Observations speak for Simplicity  $\left[ \Lambda_{UV} \gg m_{weak} \right. \left. \begin{array}{l} \mathcal{L}_{SM} \rightarrow \mathcal{L}^{d \leq 4} \quad \text{B, L, "GIM suppression", custodial symm, ...} \\ m_\nu \ll m_{weak} \quad \text{beautifully explained} \end{array} \right.$

Theory expects Naturalness  $\left[ \delta m_h^2 \sim \frac{y_t^2}{4\pi^2} \Lambda_{UV}^2 + \dots \right. \rightarrow \Lambda_{UV} \lesssim 500 \text{ GeV}$

Clash between Simplicity and Naturalness

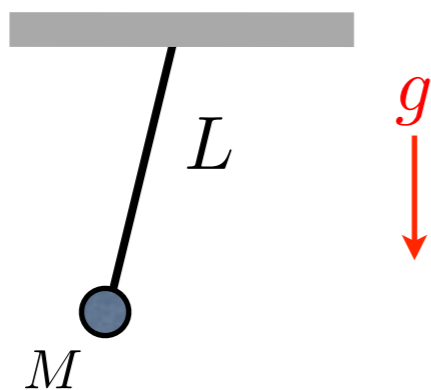
Made concrete by all available Natural models (SUSY, Comp Higgs,...)



$$m_h^2 = c \frac{y_t^2}{4\pi^2} \Lambda_{UV}^2 + \dots$$

high spin symm  $\swarrow$   $\searrow$  dilatation symm

As good as dimensional analys in mechanics



$$\omega = c \sqrt{\frac{g}{L}}$$

***Fine Tuning***: violation of expectations from symmetry and dim. analysis

$$\epsilon_T \equiv \frac{m_H^2 |_{observed}}{m_H^2 |_{expected}}$$

***Criticality***  ***Fine Tuning***  ***Landscape***

Ex. Quantum criticality in anti-ferromagnet

Sachdev '09

$$V(\vec{S}) = m^2(P) \vec{S} \cdot \vec{S} + \lambda (\vec{S} \cdot \vec{S})^2 \quad m^2(P) = m_0^2 \left( 1 - \frac{P}{P_c} \right)$$

Can undo ***natural*** expectation from atomic physics by ***tuning*** the pressure at a ***critical*** value in a ***landscape*** of possibilities

# The two Chief Systems

I. The SM is valid up to  $\Lambda_{UV} \gg TeV$

- B, L and Flavor: beautifully in accord with observation
- Higgs mass & C.C. hierarchy point beyond naturalness
  - anthropic selection
  - failure of EFT ideology (UV/IR connection)
  - ...

II. Naturalizing New Physics appears at  $\Lambda_{UV} \sim 1 TeV$

- Constraints on B, L, Flavor & CP only met by clever model building

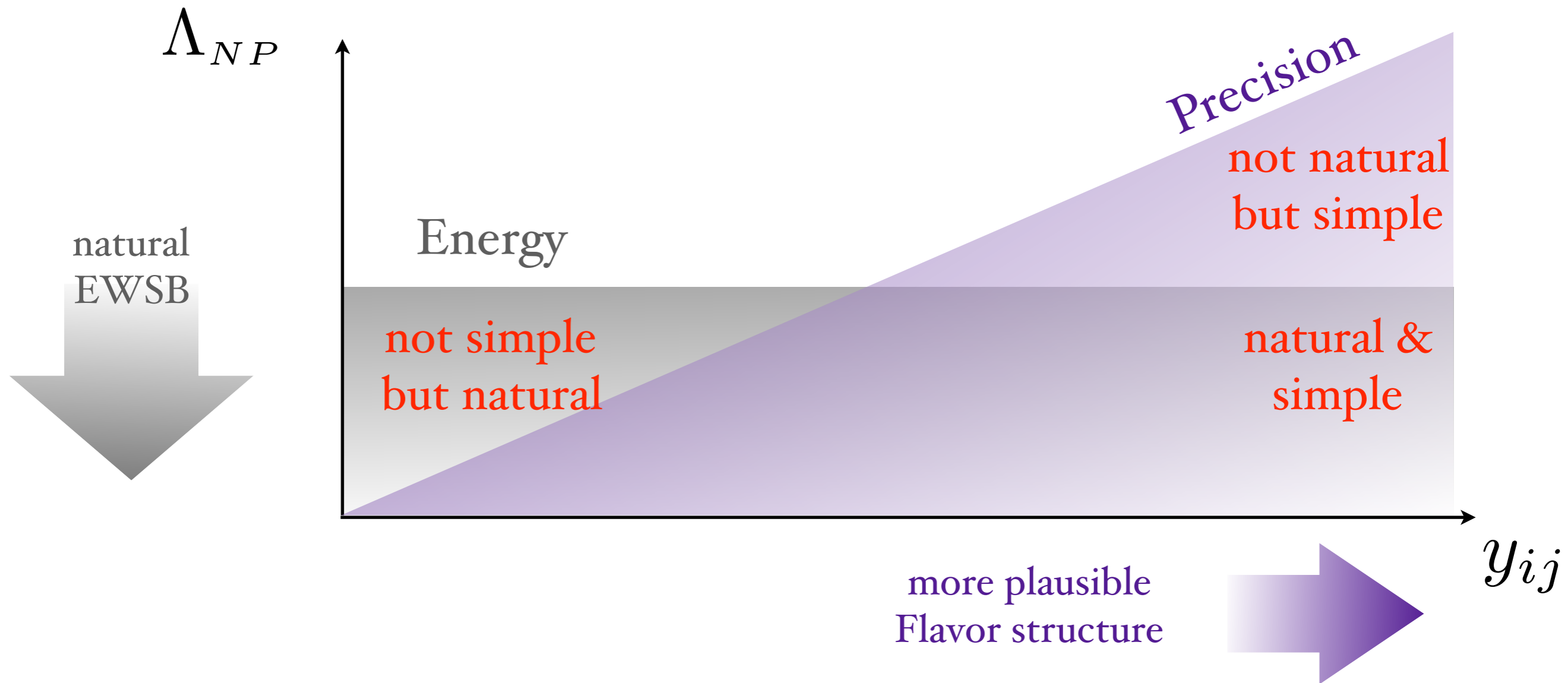
Simplicity



Naturalness

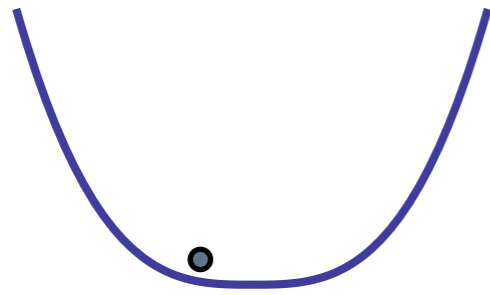
# Complementarity of Energy and Precision

$$\mathcal{L}_{eff} = \frac{y_{ijkl}}{\Lambda_{NP}^2} \bar{q}_i q_j \bar{q}_k q_l + m_i \frac{y_{ij}}{\Lambda_{NP}^2} \bar{q}_i \sigma_{\mu\nu} q_j F^{\mu\nu} + \dots$$

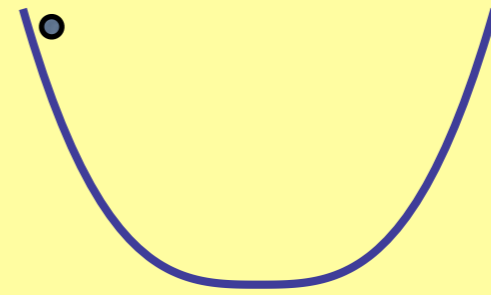


Laws

Phenomena



add quanta




go back in time

add gravity

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = T_{\mu\nu}$$

History

- FRW  isotropy and homogeneity at large scales

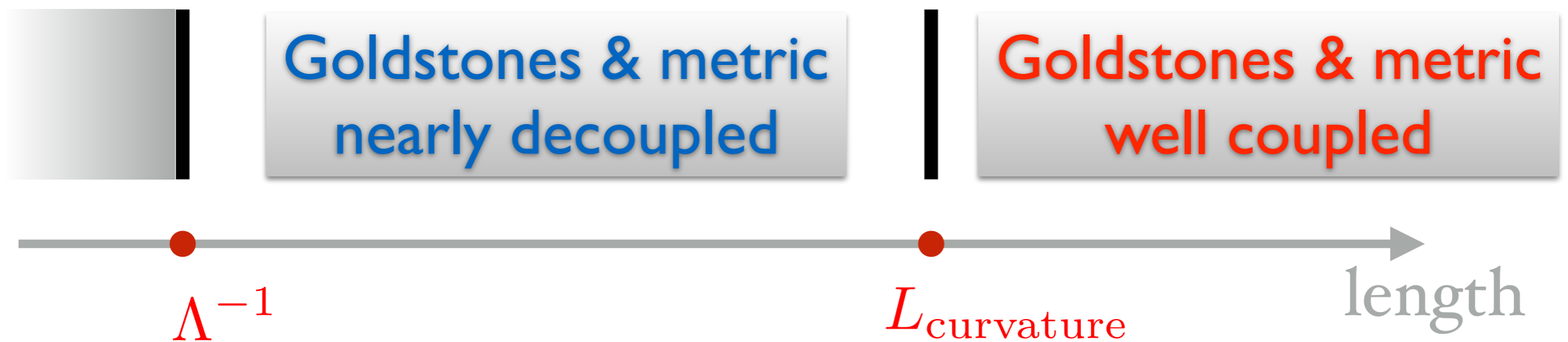
$T_{\mu\nu}$   finite density phase spontaneously breaking spacetime symmetry down to euclidean group  $ISO(3)$

- macroscopic dynamics universally described by hydrodynamics modes (Goldstone bosons)

- gravity + hydrodynamics modes

long distance  
dynamics modified

- similar to (photon + Cooper pair) in superconductor



Ex: hot plasma

$$\Lambda \sim T$$

$$L_{\text{curvature}} = H^{-1} \sim \frac{M_P}{T^2}$$

# The picture of the connubium dates back to pre-Higgs days

PHYSICAL REVIEW

VOLUME 130, NUMBER 1

1 APRIL 1963

## Plasmons, Gauge Invariance, and Mass

P. W. ANDERSON

*Bell Telephone Laboratories, Murray Hill, New Jersey*

(Received 8 November 1962)

Schwinger has pointed out that the Yang-Mills vector boson implied by associating a generalized gauge transformation with a conservation law (of baryonic charge, for instance) does not necessarily have zero mass, if a certain criterion on the vacuum fluctuations of the generalized current is satisfied. We show that the theory of plasma oscillations is a simple nonrelativistic example exhibiting all of the features of Schwinger's idea. It is also shown that Schwinger's criterion that the vector field  $m \neq 0$  implies that the matter spectrum before including the Yang-Mills interaction contains  $m = 0$ , but that the example of superconductivity illustrates that the physical spectrum need not. Some comments on the relationship between these ideas and the zero-mass difficulty in theories with broken symmetries are given.

It is noteworthy that in most of these cases, upon closer examination, the Goldstone bosons do indeed become tangled up with Yang-Mills gauge bosons and, thus, do not in any true sense really have zero mass. Superconductivity is a familiar example, but a similar phenomenon happens with phonons; when the phonon frequency is as low as the gravitational plasma frequency,  $(4\pi G\rho)^{1/2}$  (wavelength  $\sim 10^4$  km in normal matter) there is a phonon-graviton interaction: in that case, because of the peculiar sign of the gravitational interaction, leading to instability rather than finite

mass.<sup>12</sup> Utiyama<sup>13</sup> and Feynman have pointed out that gravity is also a Yang-Mills field. It is an amusing observation that the three phonons plus two gravitons are just enough components to make up the appropriate tensor particle which would be required for a finite-mass graviton.



Ex 1: relativistic  
superfluid

~~$P_0$~~ ,  ~~$K_i$~~ ,  ~~$Q$~~  broken

$$\bar{P}_0 = P_0 - \mu Q, \quad \bar{P}_i = P_i, \quad \bar{J}_i = J_i$$

$$\left[ \begin{array}{l} Q : \phi \rightarrow \phi + c \quad \Longrightarrow \quad \mathcal{L} \equiv \mathcal{L}(\partial\phi) \\ \phi = \mu t + \pi \quad \longrightarrow \quad \text{phonon} \end{array} \right.$$

- Add small explicit Q breaking:  $V(\phi)$   $V(\phi)'' \ll H^2$

$\pi$   $\left[ \begin{array}{l} \text{Goldstone of } \cancel{P_0} \\ \text{pseudo-Goldstone of } \cancel{Q} \end{array} \right.$

Effective Field  
Theory of Inflation

Fierz-Pauli massive gravity



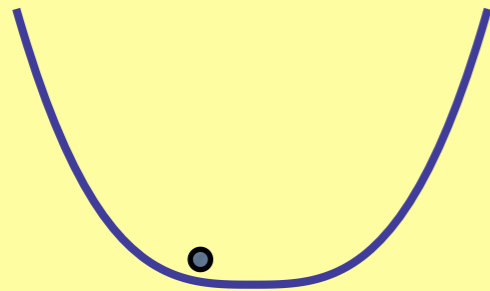
relativistic super-solid coupled to gravity

Ex 2:

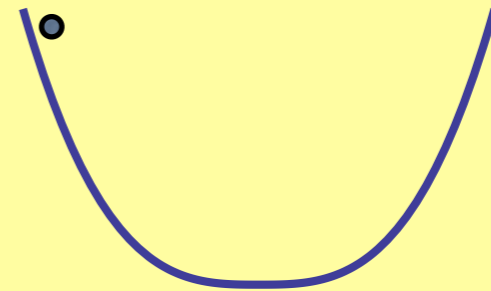
Arkani-Hamed, Georgi, Schwartz '02

Laws

Phenomena



add quanta



go back in time

add gravity

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = T_{\mu\nu}$$

History

# Feynman diagrams through superfluids

with Gil Badel, Gabriel Cuomo, Alexander Monin, [arXiv:1909.01269](https://arxiv.org/abs/1909.01269)

# [Weak vs Strong] & [Classical vs Quantum]

- ▲ Weak coupling: loop expansion around leading trajectory  $\gamma_{cl}$

$$e^{-W} = e^{-[S_0 + S_1 + S_2 + \dots]}$$

- ▲ Strong coupling: PI cannot be described by leading trajectory

Common practice: few legs in weakly coupled QFT  
= small fluctuations around trivial trajectory

However when the number of legs grows expansion breaks down  
*see old review by Rubakov, arXiv:9511236, 1995*

How do we describe physics in this regime?

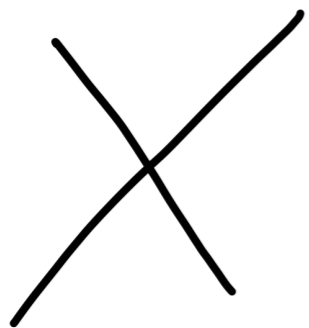
▲ Charged  $\phi^4$   $D = 4 - \epsilon$  dimension

$$\mathcal{L} = \partial_\mu \bar{\phi} \partial^\mu \phi + \frac{\lambda}{4} (\bar{\phi} \phi)^2$$

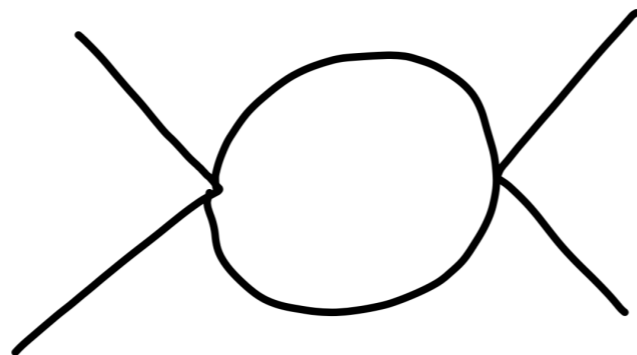
▲ Conformal invariant at Wilson-Fisher fixed point

$$\frac{\lambda_*}{(4\pi)^2} = \frac{\epsilon}{5} + \frac{3\epsilon^2}{25} + O(\epsilon^3)$$

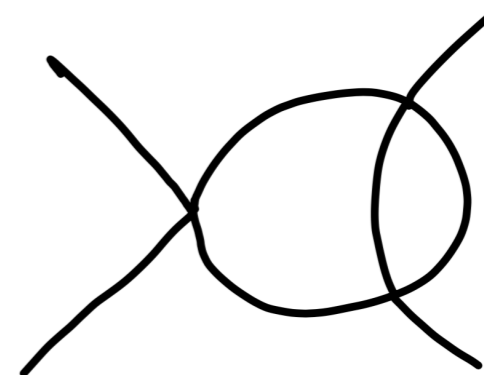
# Few Legs



$$\lambda$$

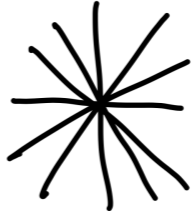


$$\frac{\lambda^2}{16\pi^2}$$

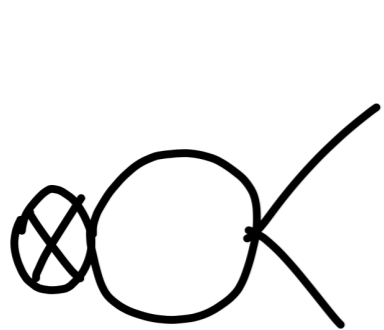


$$\frac{\lambda^3}{(16\pi^2)^2}$$

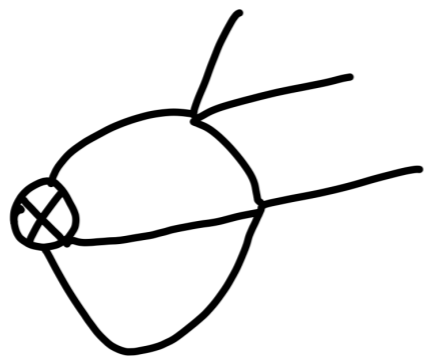


Many Legs:  $\phi^n$   $n \gg 1$    $\equiv \otimes$

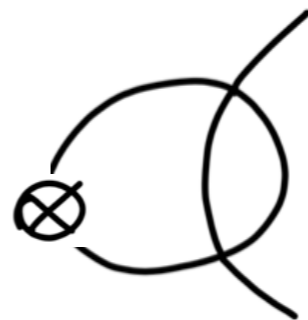
$$\langle \phi^n(x) \bar{\phi}^n(0) \rangle \propto \frac{1}{x^{2\Delta_n}} \quad \Delta_n \equiv \frac{D-2}{2} + \gamma_n$$



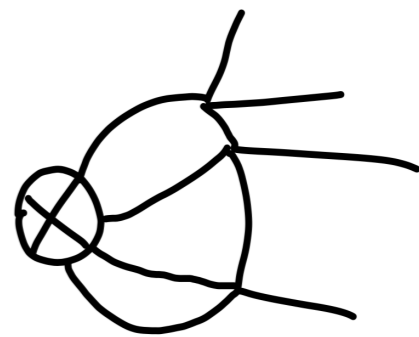
$$\lambda n(n-1)$$



$$\lambda^2 n(n-1)(n-2)$$



$$\lambda^2 n(n-1)$$



$$\lambda^3 n(n-1)(n-2)(n-3)$$

perturbation theory breaks down at  $\frac{\lambda n}{16\pi^2} \gtrsim 1$

series can be organized as a double expansion

$$\frac{\gamma_n}{n} = P_0(\lambda n) + \lambda P_1(\lambda n) + \lambda^2 P_2(\lambda n) + \dots$$

similar to RG  $F_0(\lambda \text{Log}) + \lambda F_1(\lambda \text{Log}) + \dots$

or to 't Hooft large-N expansion

$$\frac{\gamma_n}{n} = P_0(\lambda n) + \frac{1}{n} \bar{P}_1(\lambda n) + \frac{1}{n^2} \bar{P}_2(\lambda n) + \dots$$

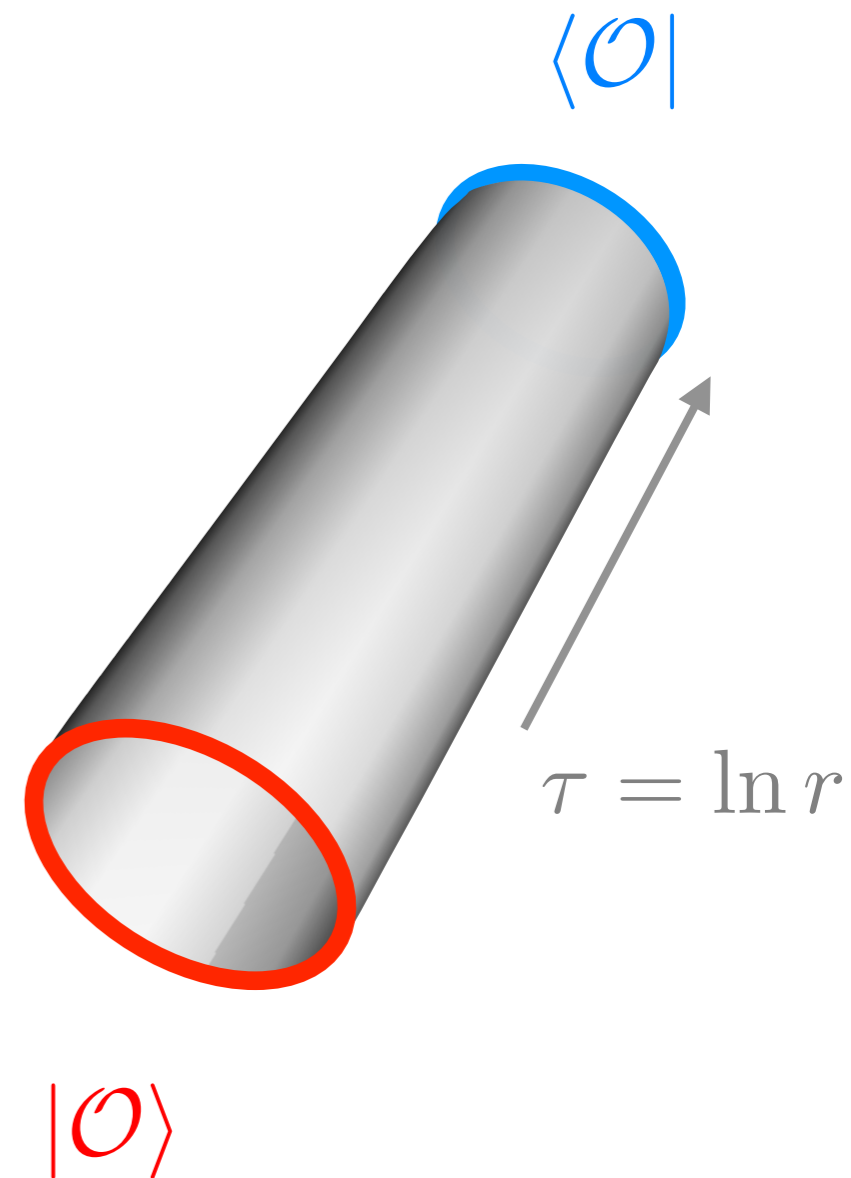
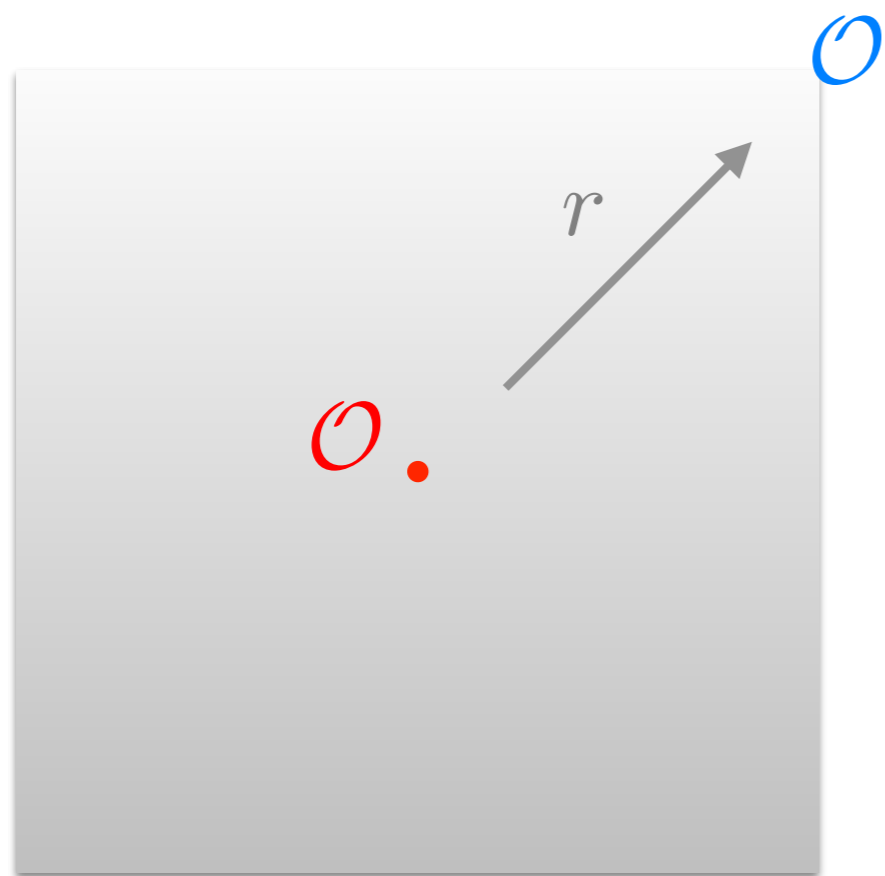
▲ What is the physics behind this?

▲ Can one compute the  $\lambda_n$  series?

Common answer:

Semiclassical expansion around non-trivial trajectory

# Mapping to the cylinder & operator/state correspondence



$$\langle \mathcal{O}(r) \mathcal{O}(0) \rangle = \frac{1}{r^{2\Delta}}$$



$$\langle \mathcal{O} | e^{-H\tau} | \mathcal{O} \rangle = e^{-\Delta\tau}$$

- path integral dominated by superfluid configuration

$$\rho = \text{const}$$

$$\phi_{cl} = \rho e^{i\chi}$$

$$\chi = -i\mu\tau$$

- plug back into action and perform systematic loop expansion around classical trajectory

$$\Delta_{\phi^n} = \frac{1}{\lambda_*} \Delta_{-1}(\lambda_* n) + \Delta_0(\lambda_* n) + \lambda_* \Delta_1(\lambda_* n) + \dots$$

# Leading order

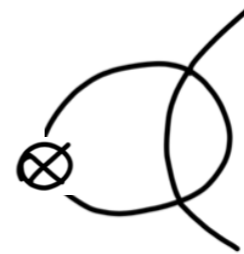
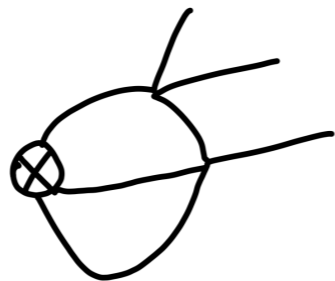
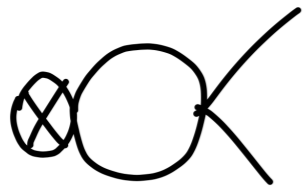
$$\frac{1}{\lambda_* n} \Delta_{-1} = \frac{3 \left[ 9x - \sqrt{81x^2 - 3} \right]^{1/3} + 3^{2/3} \left[ 9x - \sqrt{81x^2 - 3} \right]}{\left[ \left( 9x - \sqrt{81x^2 - 3} \right)^{2/3} + 3^{1/3} \right]^2} + \frac{9 \times 3^{1/3} x \left[ 9x - \sqrt{81x^2 - 3} \right]^{2/3}}{2 \left[ \left( 9x - \sqrt{81x^2 - 3} \right)^{2/3} + 3^{1/3} \right]^2}$$
$$x \equiv \frac{\lambda_* n}{16\pi^2}$$

Supposed to resum leading powers of  $n$  at all loops!

expanding at small  $\lambda n$

$$\Delta_{\phi^n} = n + \frac{\lambda n^2}{32\pi^2} - \frac{\lambda^2 n^3}{512\pi^4} + \frac{\lambda^3 n^4}{4096\pi^6} + O(\lambda^4 n^5)$$

and comparing with diagrams



$$\gamma_n = \frac{\lambda n(n-1)}{32\pi^2} - \frac{\lambda^2 n^2(n-1)}{512\pi^4} + \dots \quad \text{they happily agree}$$

▲  $1/n$  suppressed terms  $\longleftrightarrow$  Casimir energy of superfluid

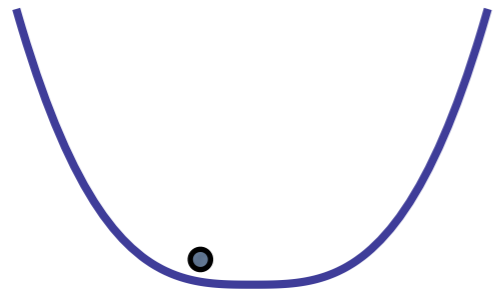
▲  $\epsilon \rightarrow 1$  extrapolation well matches Monte Carlo simulations of U(1) model in D=3

▲ spectrum of 'nearby' operators described by phonon spectrum

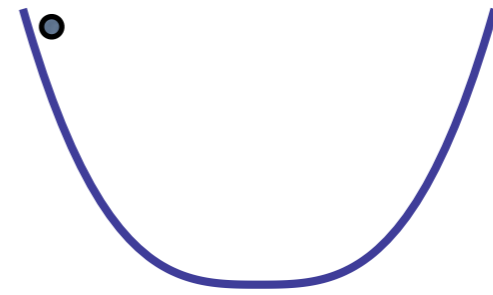
$\phi^{n-2} \partial_\mu \phi \partial_\nu \phi$   $\longleftrightarrow$  phonon with  $\ell = 2$



Laws



Phenomena



$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = T_{\mu\nu}$$

History