

Studying Gravitational Waves with the PENCIL CODE

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PENCIL CODE as a GW solver

PENCIL CODE special modules for GWs:

- *gravitational_waves.f90*;
- *gravitational_waves_hij6.f90*;
- ✓ *gravitational_waves_hTXk.f90*.

Coupled to HD/MHD source module:

- *hydro.f90*
- *hydro_kinematic.f90*
- *magnetic.f90*
- *magnetic/maxwell.f90*
- ...

or another special modules (?)

Studying GWs using the PENCIL CODE

- *The timestep constraint in solving the gravitational wave equations sourced by hydromagnetic turbulence* [Roper Pol, Brandenburg, et al. 2020]
- *Numerical simulations of gravitational waves from [early-universe](#) turbulence* [Roper Pol, Mandal, et al. 2020]
- *[Circular polarization](#) of gravitational waves from early-Universe helical turbulence* [Kahniashvili et al. 2021]
- *Relic gravitational waves from the [chiral magnetic effect](#)* [Brandenburg, He, et al. 2021]
- *Can we observe the [QCD phase transition](#)-generated gravitational waves through pulsar timing arrays?* [Brandenburg, Clarke, et al. 2021]
- *Spectrum of turbulence-sourced gravitational waves as a constraint on [graviton mass](#)* [He, Brandenburg, and Sinha 2021]
- *The [scalar, vector, and tensor modes](#) in gravitational wave turbulence simulations* [Brandenburg, Gogoberidze, et al. 2021]

The gravitational wave (GW) equation

GW equation in terms of comoving \mathbf{x} , conformal t and scaled source T_{ij} in real space is¹

$$(\partial_t^2 - \nabla^2)h_{ij}(\mathbf{x}, t) = \mathcal{G}(t)T_{ij}(\mathbf{x}, t), \quad (1)$$

where $\mathcal{G}(t) = 6/t$, and in \mathbf{k} space is

$$\ddot{\tilde{h}}_{ij}(\mathbf{k}, t) + k^2 \tilde{h}_{ij}(\mathbf{k}, t) = \mathcal{G}(t)\tilde{T}_{ij}(\mathbf{k}, t). \quad (2)$$

Can evolve the wave equation as

$$\begin{pmatrix} k\tilde{h} - k^{-1}\mathcal{G}\tilde{T} \\ \dot{\tilde{h}} \end{pmatrix}_{+, \times}^{t+\delta t} = \begin{pmatrix} \cos k\delta t & \sin k\delta t \\ -\sin k\delta t & \cos k\delta t \end{pmatrix} \begin{pmatrix} k\tilde{h} - k^{-1}\mathcal{G}\tilde{T} \\ \dot{\tilde{h}} \end{pmatrix}_{+, \times}^t. \quad (3)$$

¹Roper Pol, Brandenburg, et al. 2020.

- Why?
 - Phase transitions during the early radiation era \Rightarrow inevitable turbulence.
- How?
 - Construct stress-energy tensor \tilde{T}_{ij} for GWs

$$\tilde{T}_{ij} = \frac{4}{3}\gamma^2\rho u_i u_j - B_i B_j + \dots \quad (4)$$

from turbulent \mathbf{u} , \mathbf{B} fields.

Turbulent sources: an example

The continuity, velocity and magnetic induction equations:

$$\frac{\partial \ln \rho}{\partial t} = -\frac{4}{3}(\nabla \cdot \mathbf{u} + \mathbf{u} \cdot \nabla \ln \rho) + \frac{1}{\rho}[\mathbf{u} \cdot (\mathbf{J} \cdot \mathbf{B}) + \eta j^2] \quad (5)$$

$$\frac{D\mathbf{u}}{Dt} = \frac{\mathbf{u}}{3}(\nabla \cdot \mathbf{u} + \mathbf{u} \cdot \nabla \ln \rho) - \frac{\mathbf{u}}{\rho}[\mathbf{u} \cdot (\mathbf{J} \times \mathbf{B}) + \eta j^2] - \frac{1}{4}\nabla \ln \rho + \frac{3}{4\rho}\mathbf{J} \times \mathbf{B} + \frac{2}{\rho}\nabla \cdot (\rho\nu\mathbf{S}) \quad (6)$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{B} - \eta \mathbf{J} + \mathcal{F}), \quad (7)$$

where $\mathbf{J} = \nabla \times \mathbf{B}$ and $D/Dt = \partial/\partial t + \mathbf{u} \cdot \nabla$. The rate-of-strain tensor elements are $S_{ij} = \frac{1}{2}(u_{i,j} + u_{j,i}) - \frac{1}{3}\delta_{ij}u_{k,k}$. Also, relativistic EOS $p = \rho/3$ has been used.

Turbulent sources: an example

The forcing to model generation of \mathbf{B} :

$$\mathcal{F}(\mathbf{x}, t) = \text{Re}[\mathcal{N}\tilde{\mathbf{f}}(\mathbf{k}) \exp(i\mathbf{k} \cdot \mathbf{x} + i\varphi)]. \quad (8)$$

Each time step t :

- Random \mathbf{k} and φ .
- $\tilde{\mathbf{f}}(\mathbf{k}) = (\mathbf{k} \times \mathbf{e}) / \sqrt{k^2 - (\mathbf{k} \cdot \mathbf{e})^2}$, where $\mathbf{e} \nparallel \mathbf{k}$.
- Normalisation $\mathcal{N} = f_0 / \delta t^{1/2}$, with f_0 varied by hand to achieve approximately the same $\mathcal{E}_B^{\max} = \langle \mathbf{B}^2 \rangle / 2$ for all runs (later).

An application: massive GWs

- Modified action²

$$S = \int d^4x (\mathcal{L}_{\text{EH}} + \mathcal{L}_{m_g} + \mathcal{L}_{\text{mat}}), \text{ where } \mathcal{L}_{m_g} = \frac{1}{4} m_g^2 \sqrt{-g} \left(h_{\mu\nu}^2 - \frac{1}{2} h^2 \right). \quad (9)$$

- Modified GW equation in k space

$$\ddot{\tilde{h}}_{+,\times}(\mathbf{k}, t) + (\mathbf{k}^2 + m_g^2) \tilde{h}_{+,\times}(\mathbf{k}, t) = \mathcal{G}(t) \tilde{T}_{+,\times}(\mathbf{k}, t). \quad (10)$$

- Nonlinear dispersion:

$$\omega = \sqrt{\mathbf{k}^2 + \omega_{\text{cut}}^2}, \quad (11)$$

where $\omega_{\text{cut}} = m_g c^2 / \hbar$ is an effective graviton mass term³.

²Fierz and Pauli 1939.

³Lee 2013.

An application: massive GWs

Spatial spectrum:

- $0.3 \leq \omega_{\text{cut}} \leq 10$;
- $2.2 \times 10^{-23} \text{eV} \lesssim m_g \lesssim 7.4 \times 10^{-22} \text{eV}$;
- $\Omega_{\text{GW}} \sim k^1 \Rightarrow k^3$;
- $h_c \sim k^{-1/2} \Rightarrow k^{3/2}$;
- Consistent change for different k_f .

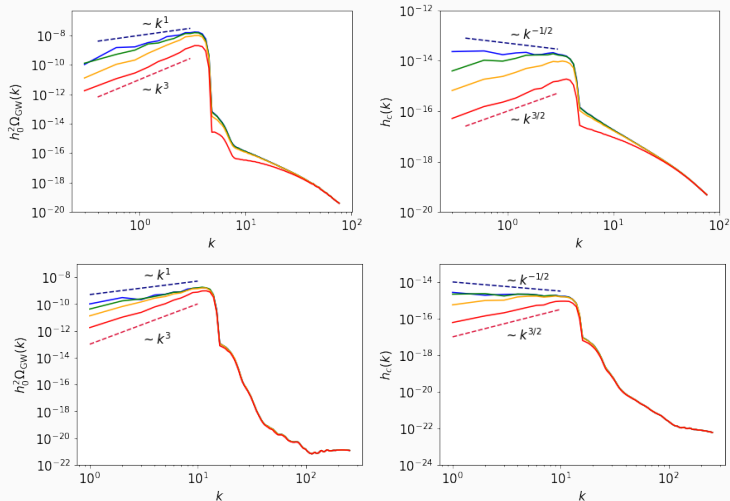


Figure 1: Spatial energy & strain spectra. $k_f = 2$ (up) & $k_f = 6$ (down).

An application: massive GWs

Temporal spectrum:

- $\omega_{\text{cut}} \in \{0, 1\}$,
- $m_g \approx 7.4 \times 10^{-23} \text{ eV}$;
- Clear cut-off;
- $f \sim 10 \text{ nHz}$;
- $\Delta t_{\text{phys}} \sim \text{yr}$.

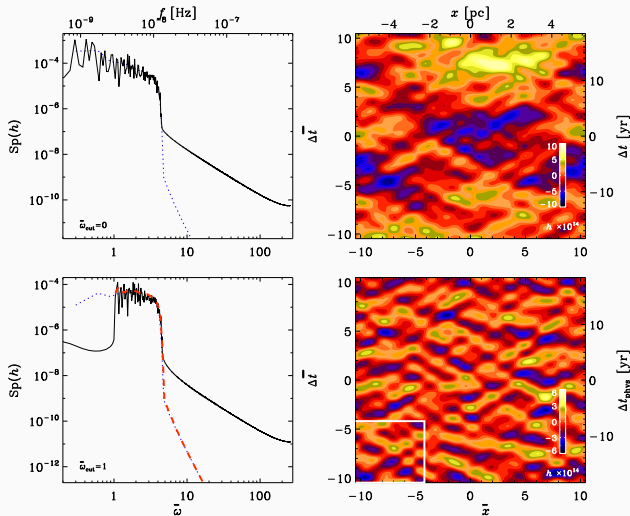


Figure 2: Temporal strain spectra. $\omega_{\text{cut}} = 0$ (up) and $\omega_{\text{cut}} = 1$ (down). 9

PENCIL CODE for GW spectral analysis:

- Turbulent source: kinetic, (electro)magnetic...
- Early universe: EWPT, QCDPT...
- Modified gravity: massive gravity and more.

Questions?