

# The Euclid Strong Lensing Pipeline

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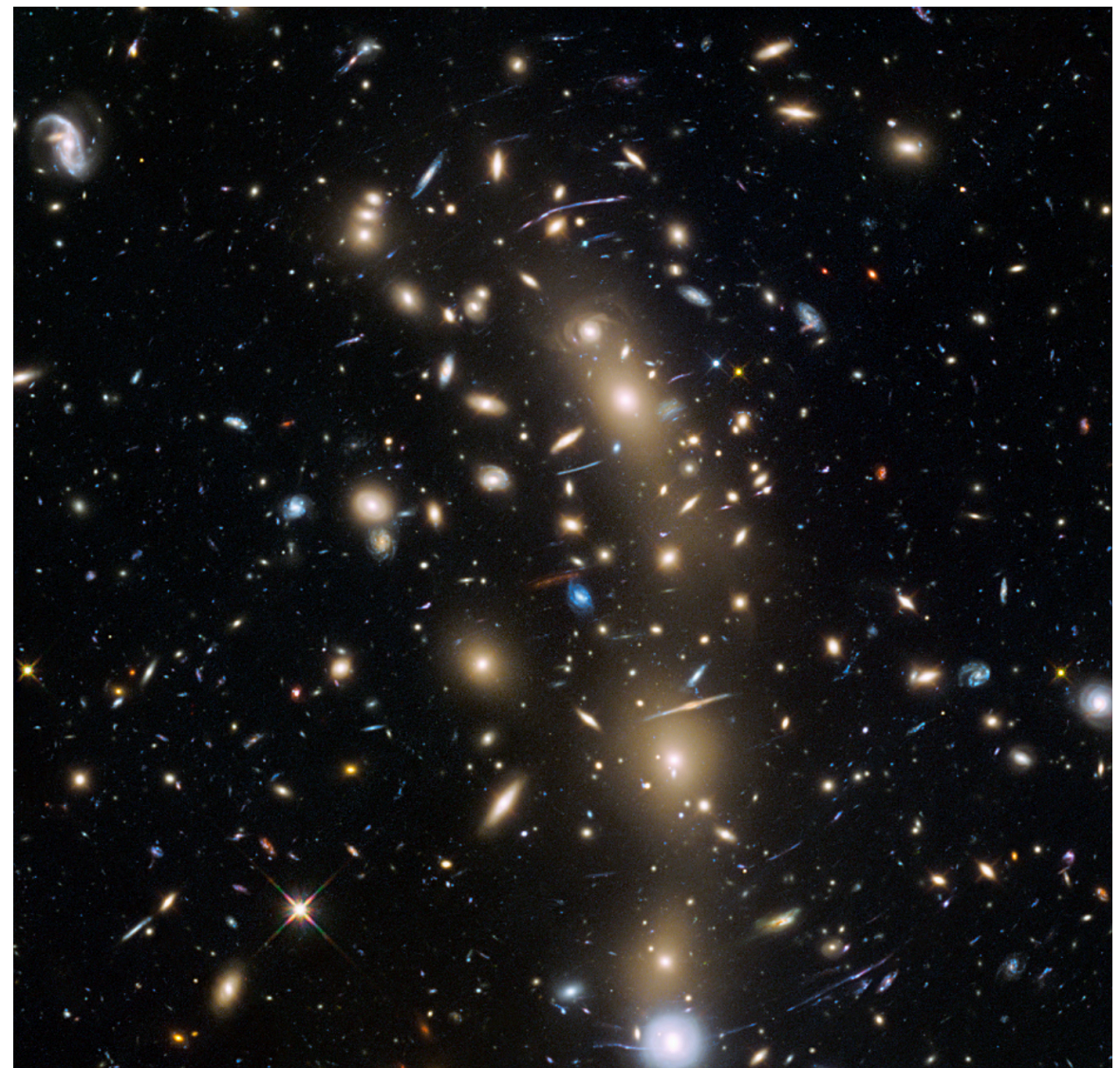
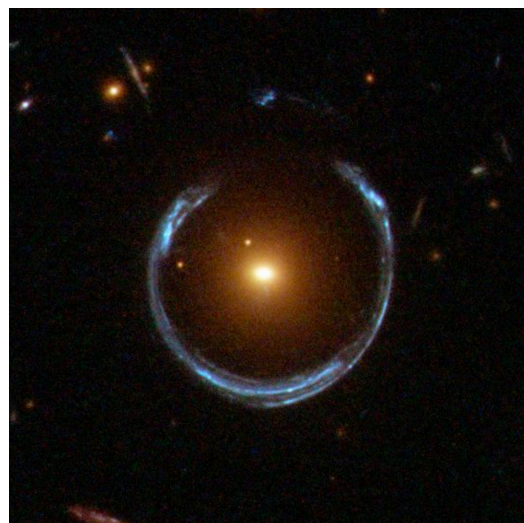
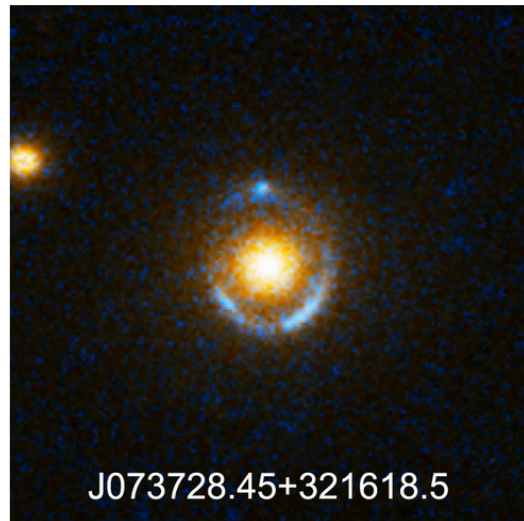
Benjamin Clément - Hubert Degaudenzi - Frédéric Courbin  
Karina Rojas - Elodie Savary  
EPFL / Unige

Euclid Swiss Days - February 2020

# Galaxy and group/cluster scales

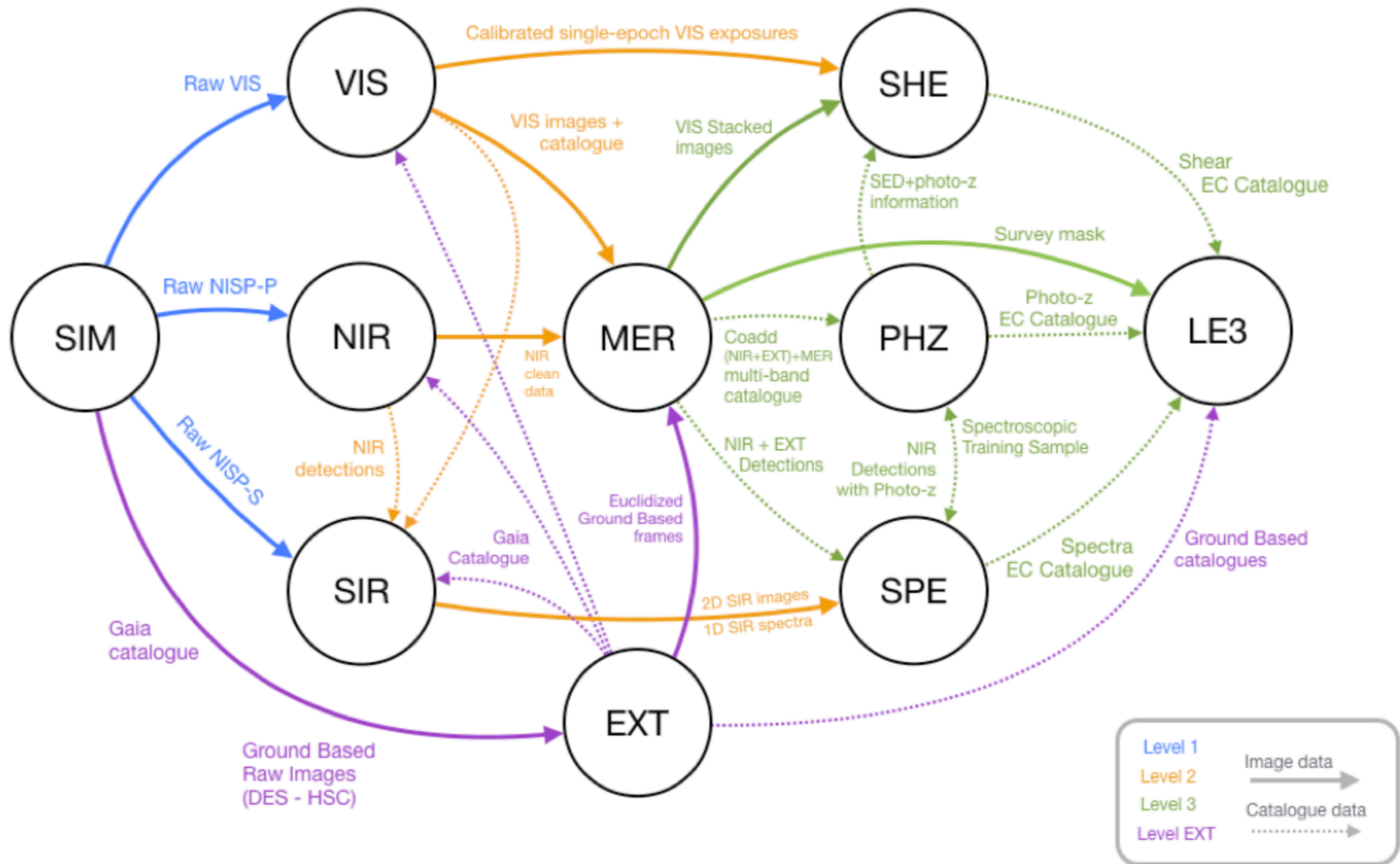
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- GGL : Galaxy-Galaxy Lens  $Re \sim 1'' - 3''$
- GrGL : Group-Galaxy Lens  $Re \sim 3'' - 10''$
- CL : Cluster Lens  $Re \sim 20'' - 50''$





# The Full Picture



# Data

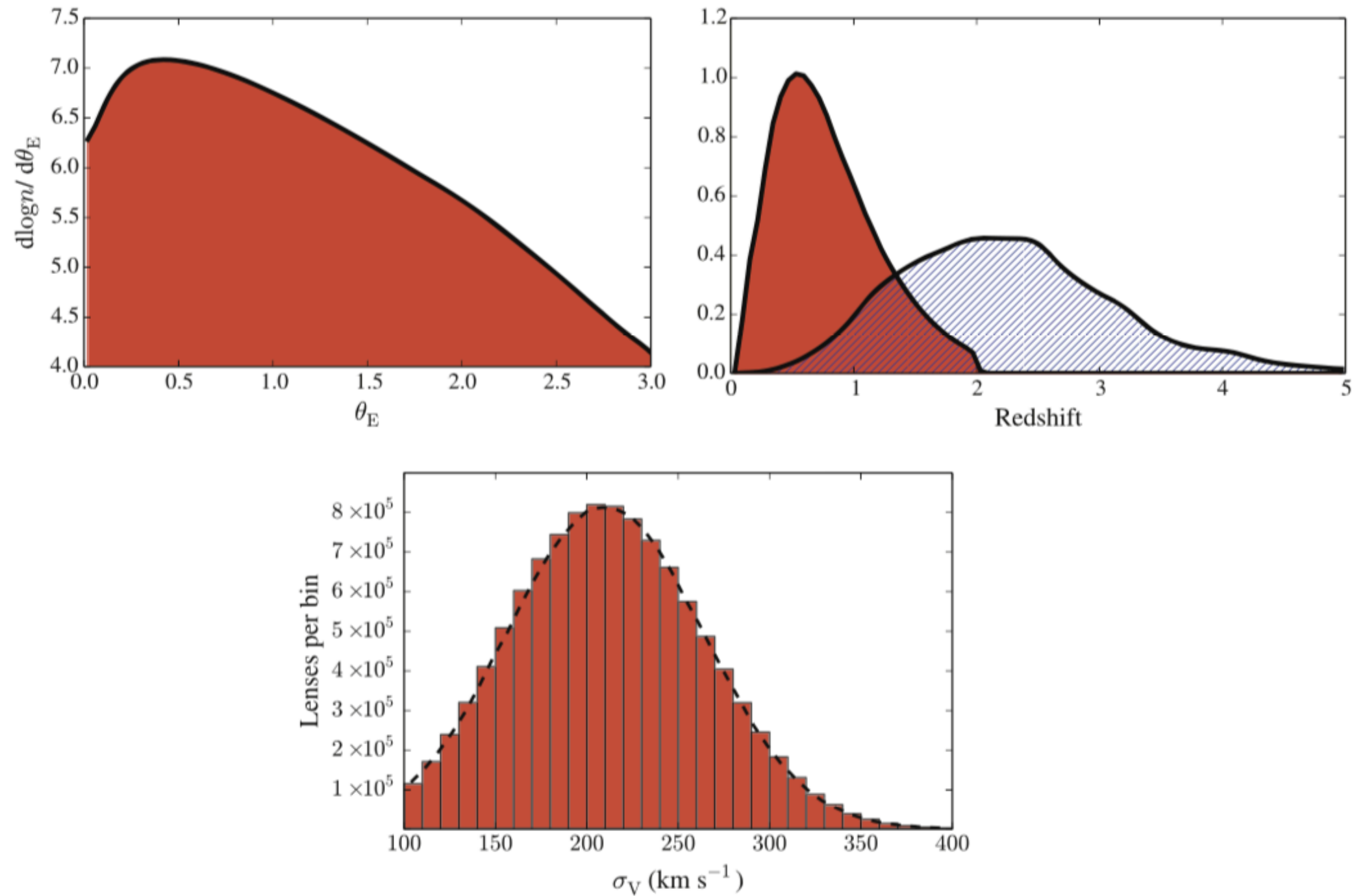
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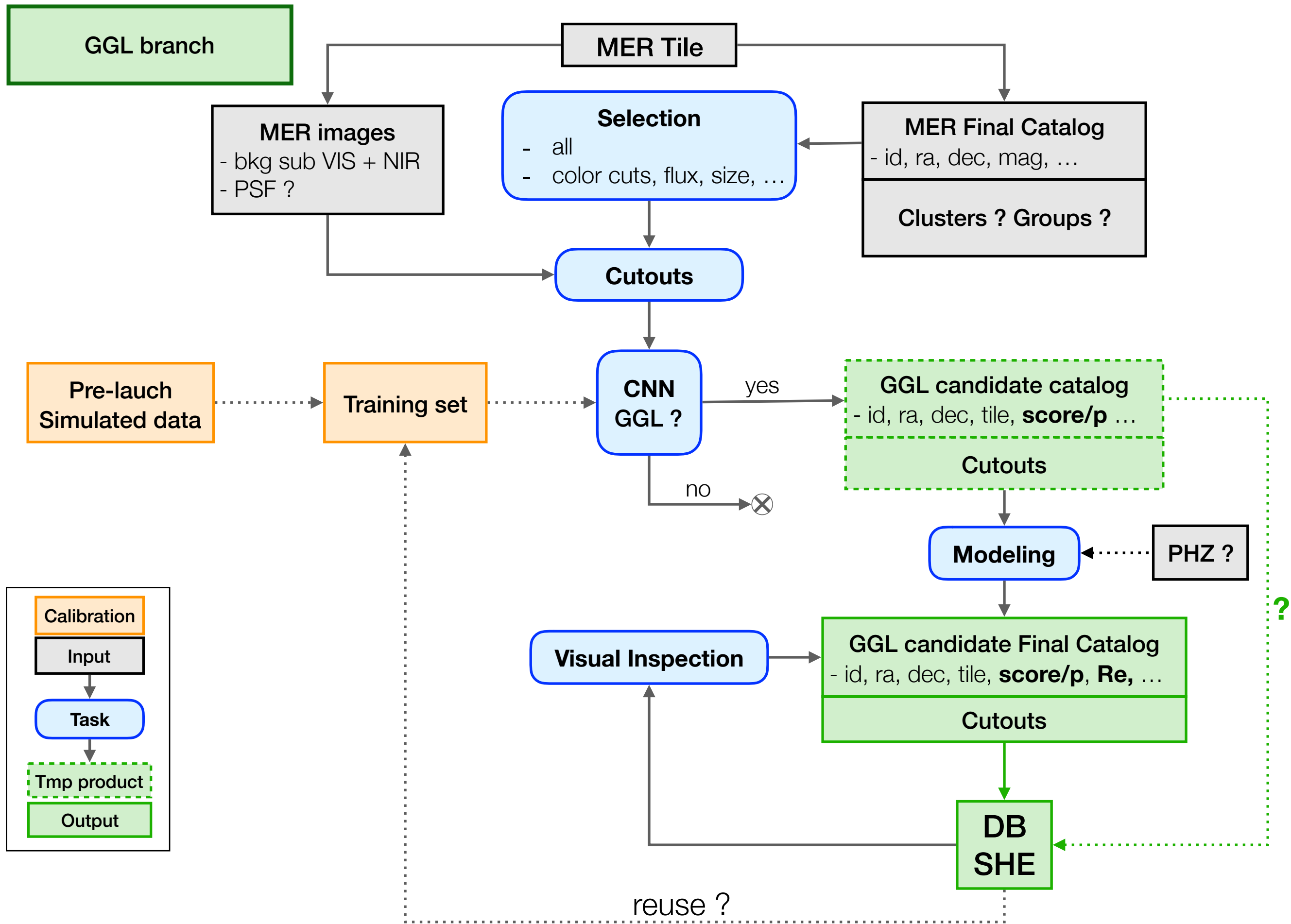
- OU-MER Tiles FOV is **32' x 32'** or **1024 arcmin<sup>2</sup>** or **0.2844 deg<sup>2</sup>**
- How many tiles in total ?  
15000 deg<sup>2</sup> ~ **50000 tiles**
- **12 bands** : **VIS, NIR (YJH)**, EXT DECAM (griz), EXT OMEGACAM (ugri)
- How many objects to inspect ?  
naive magnitude cut in VIS < 22 : ~1000 objects per tile
- How many lenses per tile to expect ?  
~170 000 lenses in the full survey (Collett 2015)  
~ 1-10 / deg<sup>2</sup>  
~ **3 lenses / tile**  
~ 60 lenses / day



# Expected properties of the lens sample

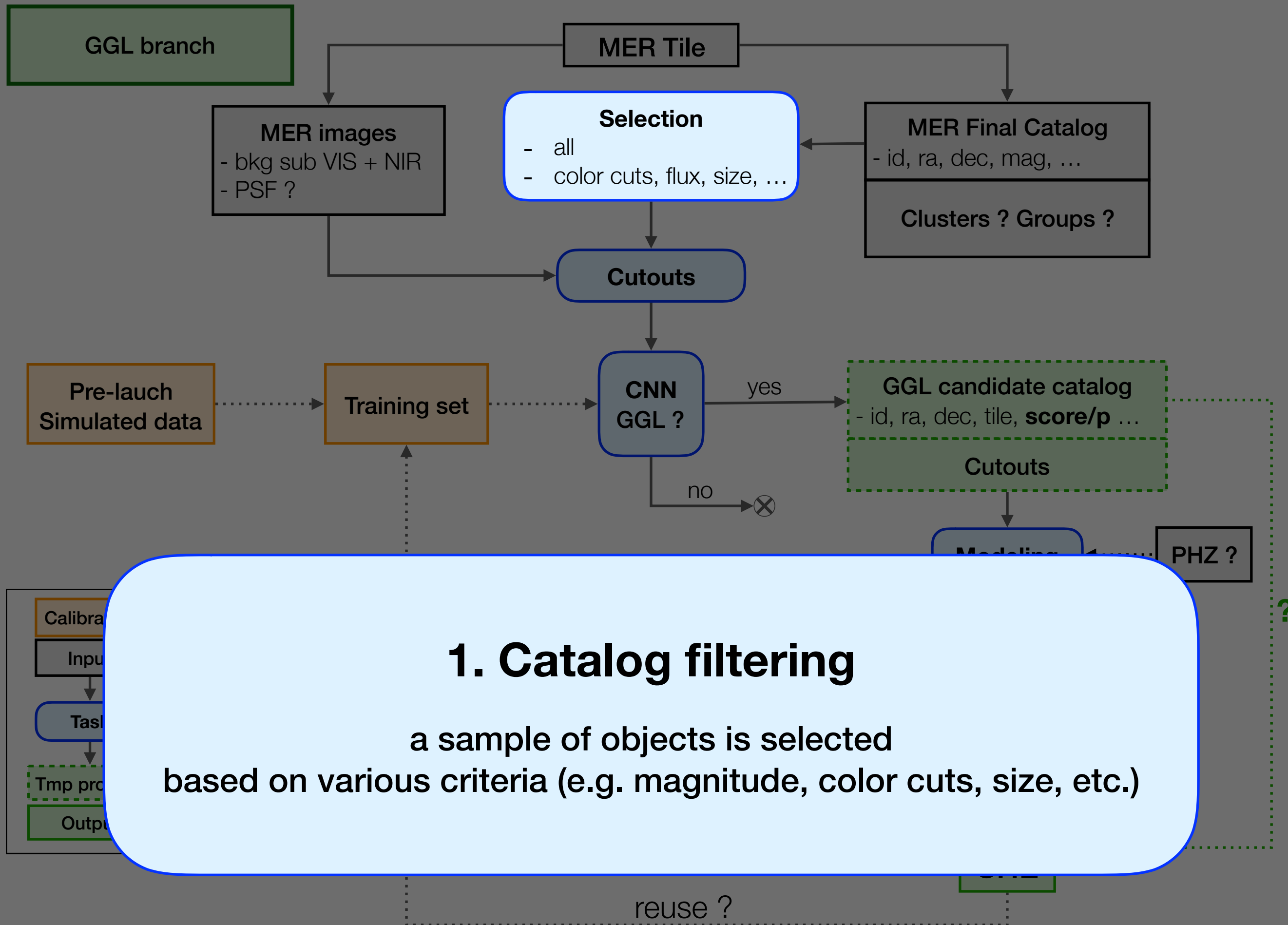
from Collett (2015)

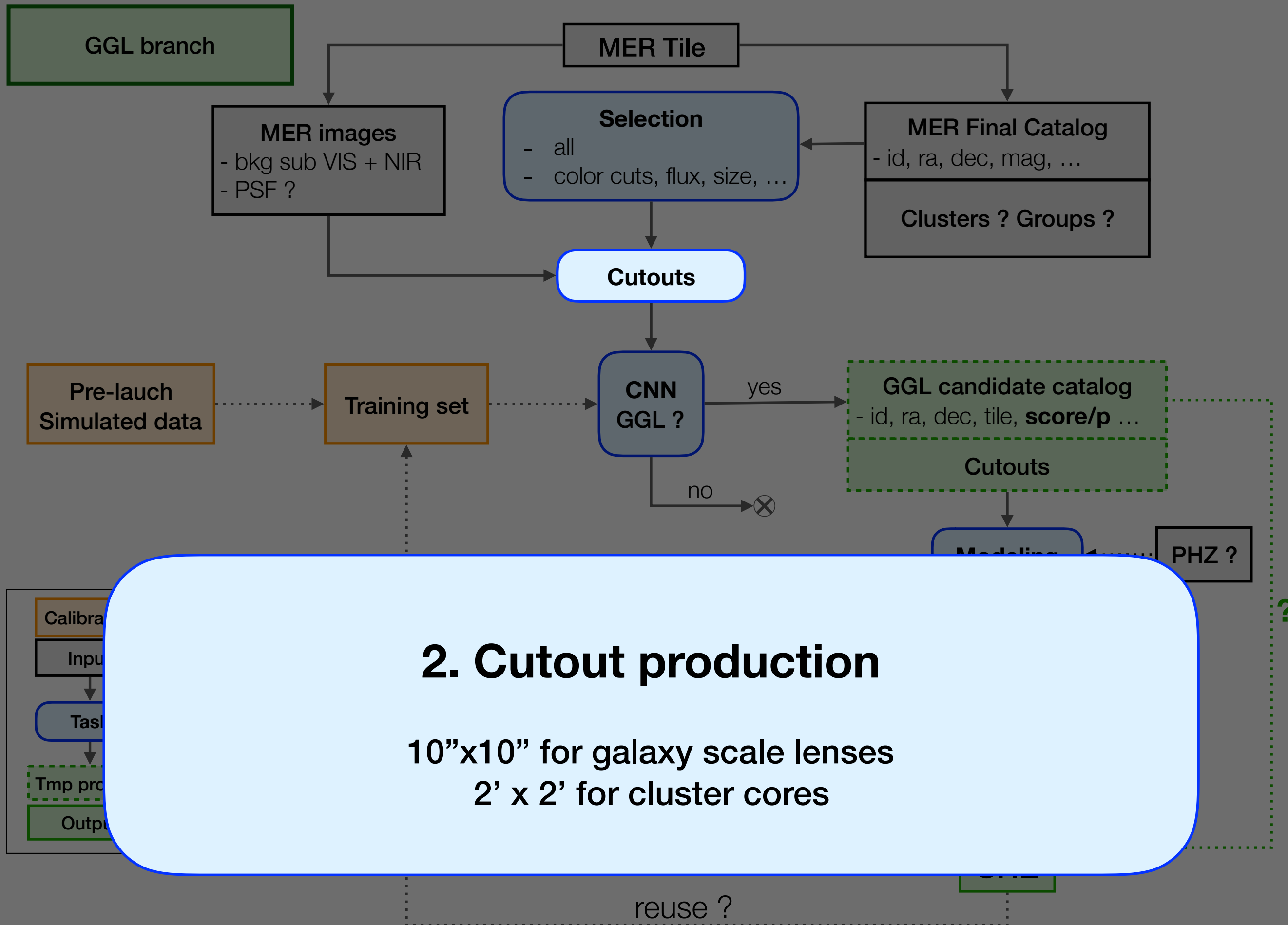


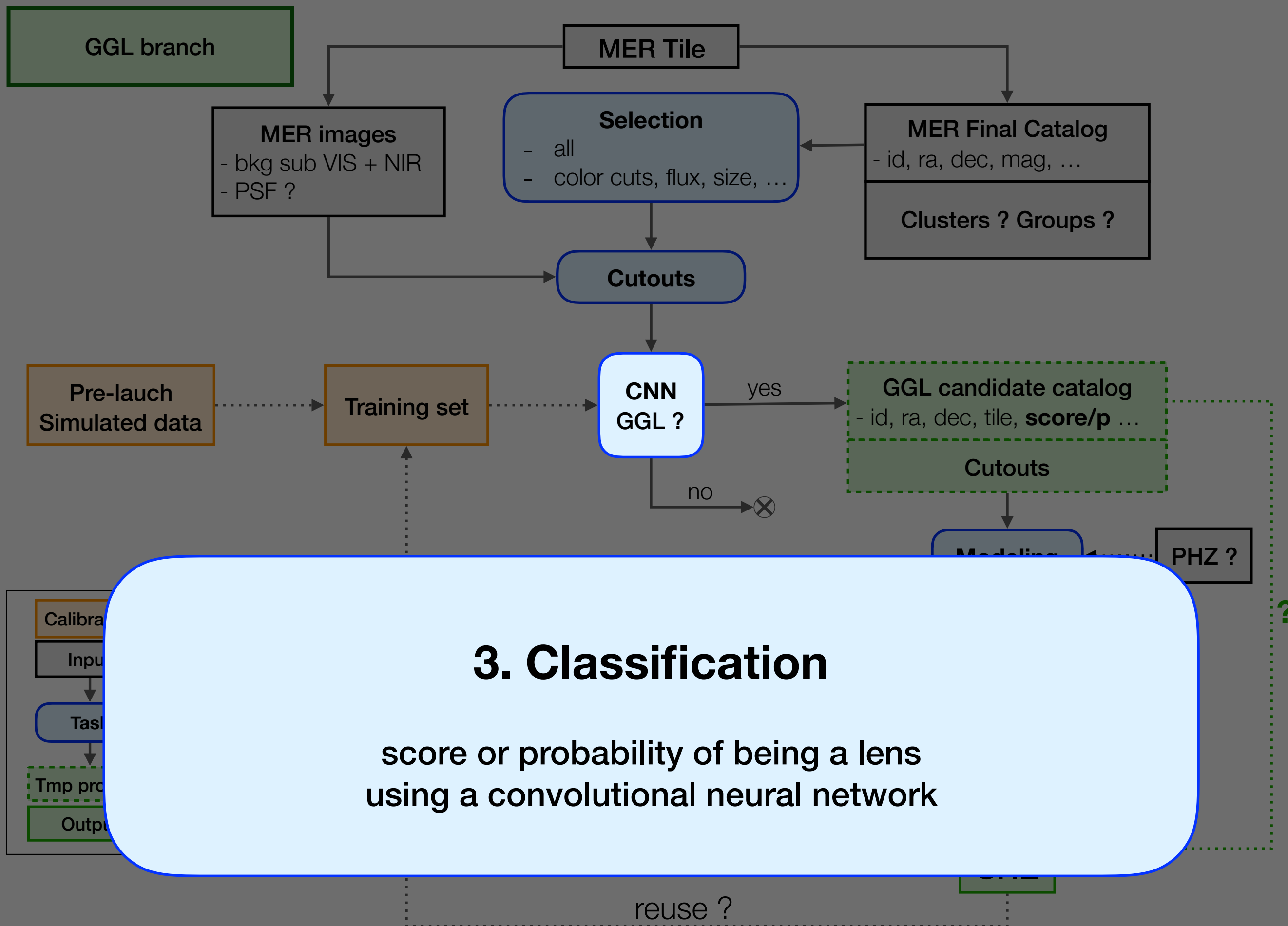


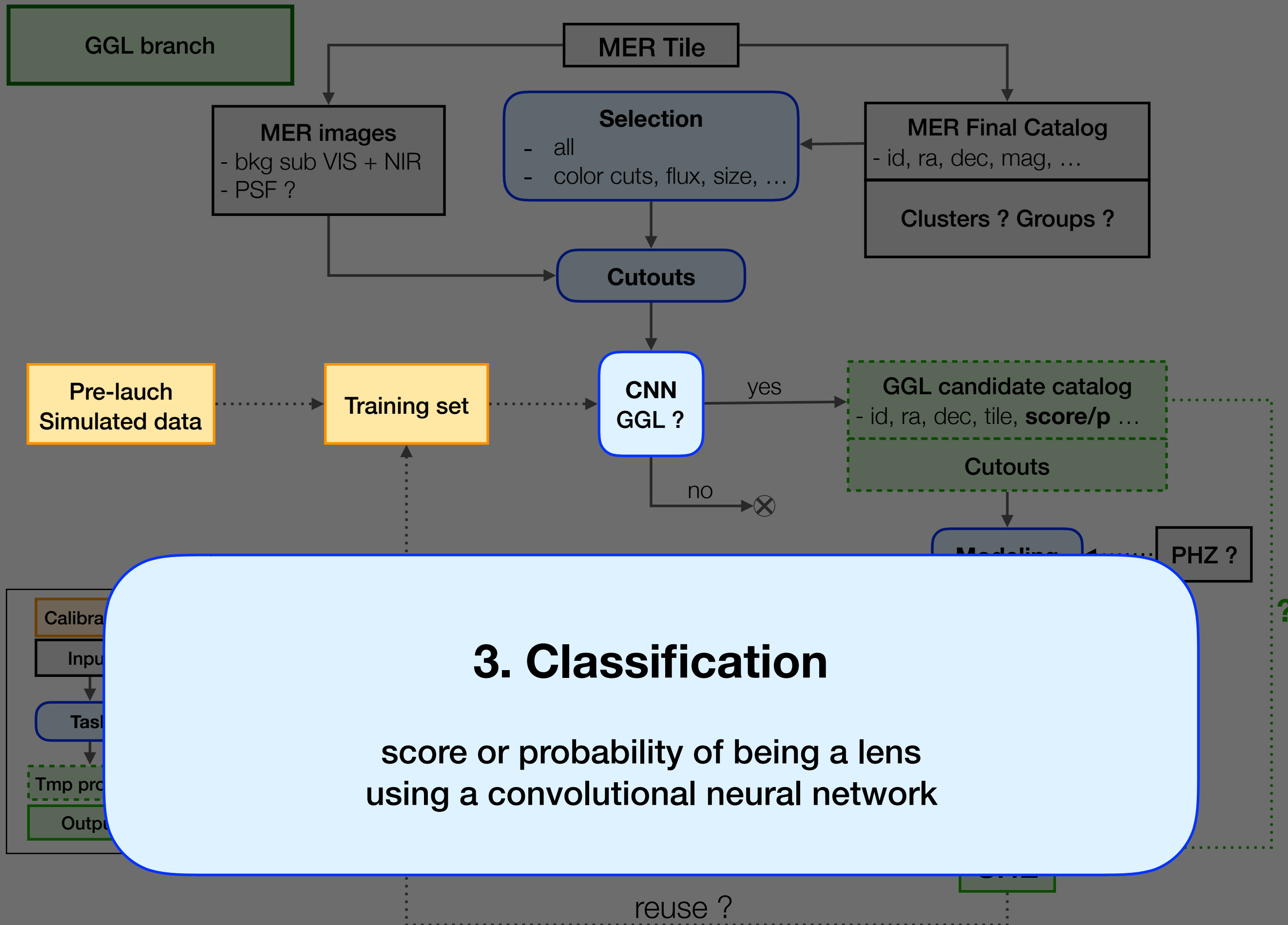




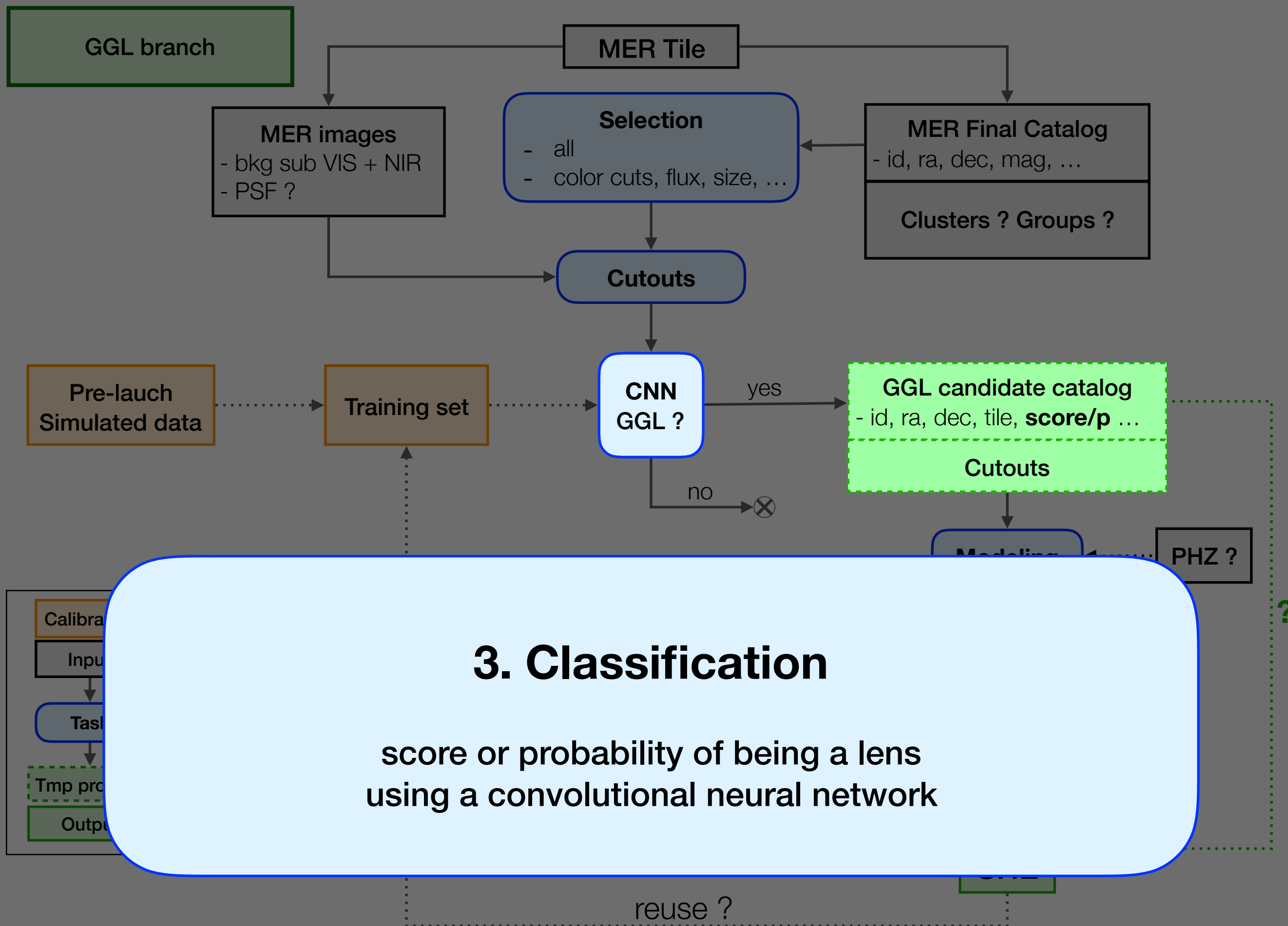












GGL branch

MER Tile

## 4. Modelling

basic physical properties :  
Einstein radius  $R_e$ , mass enclosed in  $R_e$ ,  
number of images, type of lens, ...

Pre-lens  
Simulated data

Training set

GGL ?

no

- id, ra, dec, tile, **score/p** ...

Cutouts

Modeling

PHZ ?

Visual Inspection

GGL candidate Final Catalog  
- id, ra, dec, tile, **score/p**,  $R_e$ , ...

Cutouts

DB  
SHE

reuse ?

Calibration

Input

Task

Tmp product

Output

GGL branch

MER Tile

## 5. Visual inspection

lens candidates with probability above a certain threshold are reviewed by human eye and given a grade (e.g. citizen science)

Pre-lens  
Simulated data

Training set

GGL ?

no

- id, ra, dec, tile, **score/p** ...

Cutouts

Modeling

PHZ ?

Visual Inspection

GGL candidate Final Catalog  
- id, ra, dec, tile, **score/p**, Re, ...

Cutouts

DB  
SHE

reuse ?

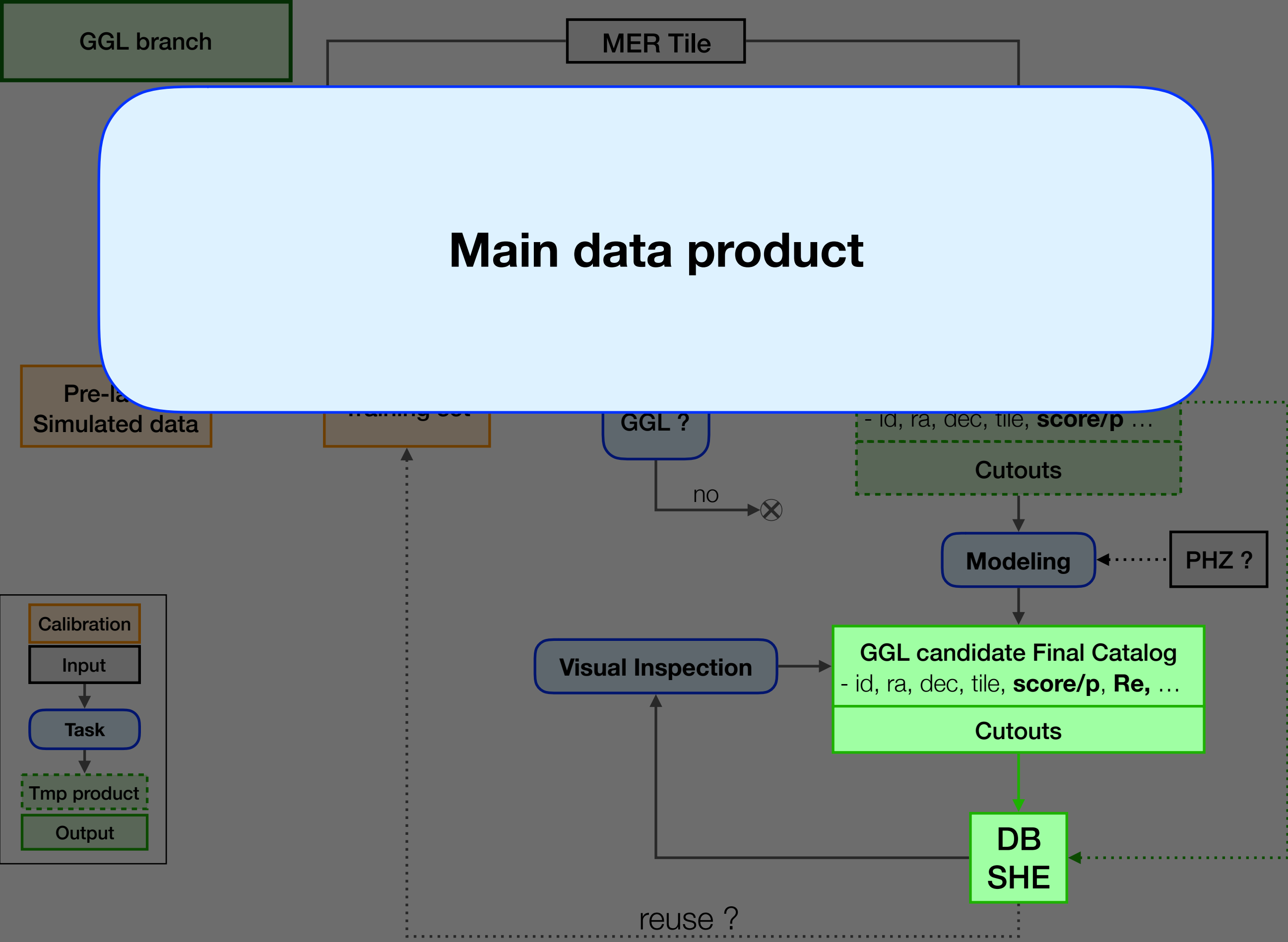
Calibration

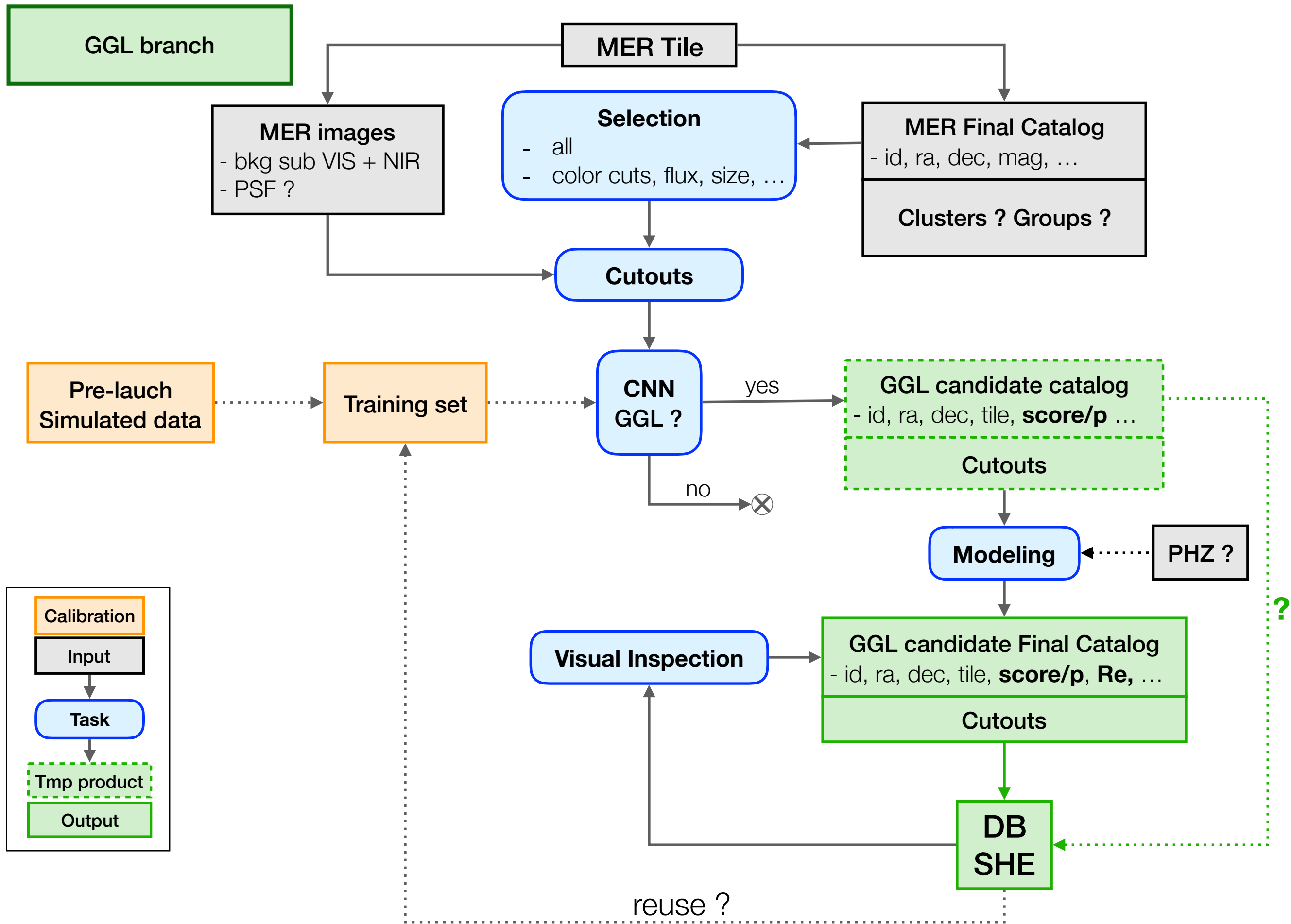
Input

Task

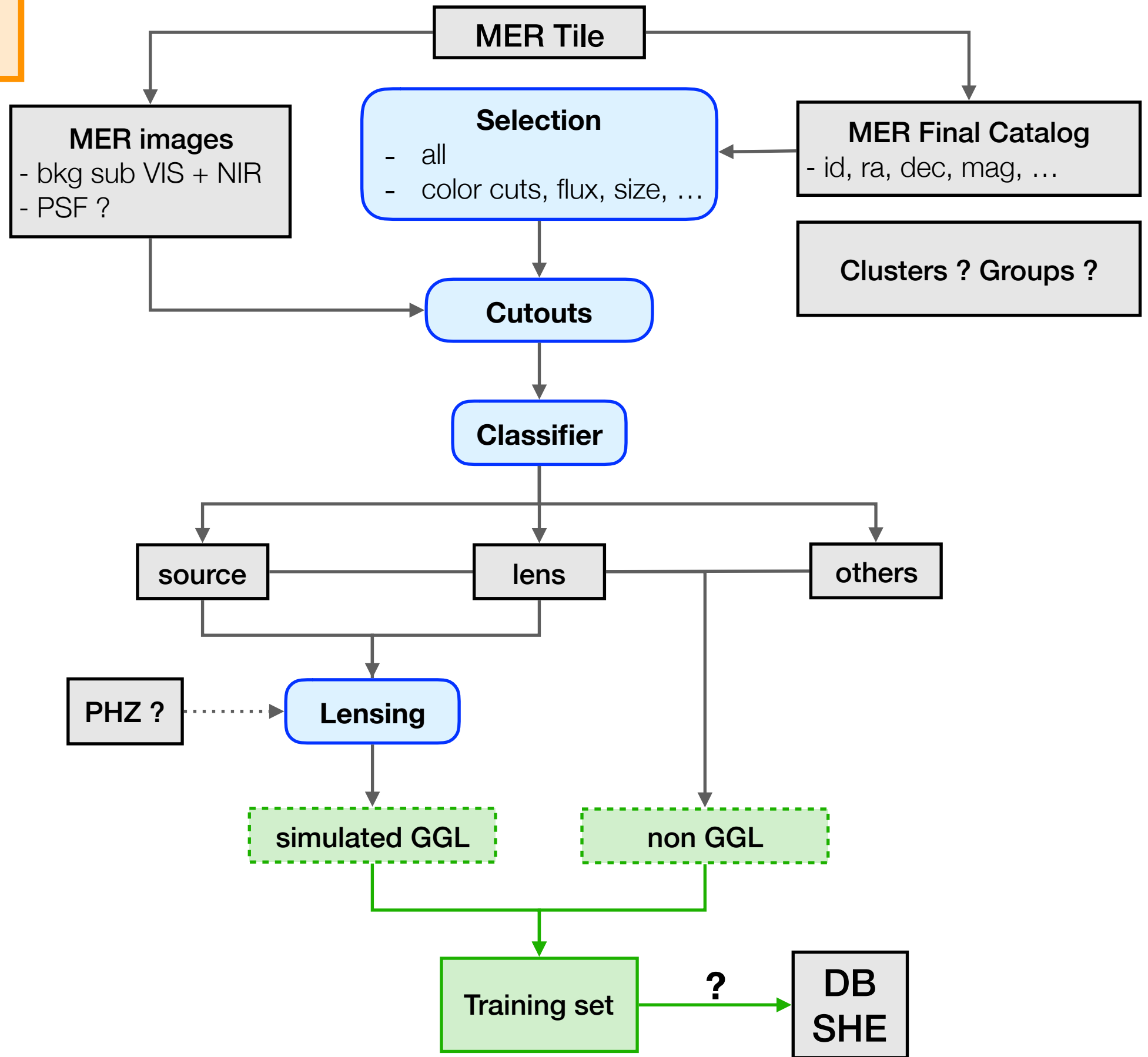
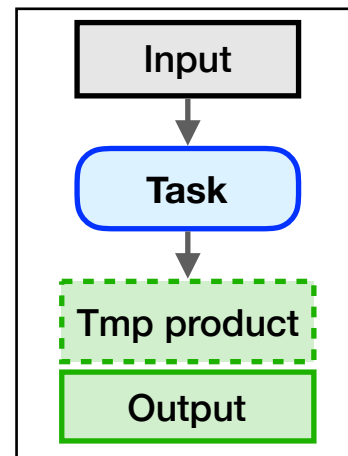
Tmp product

Output





Training set branch

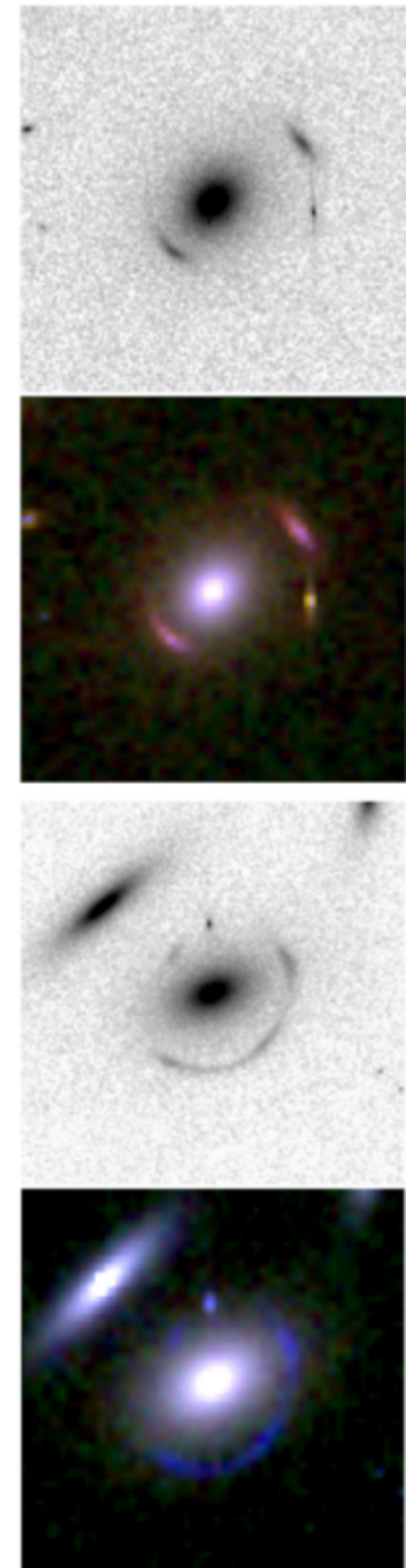




# Simulating the lens sample

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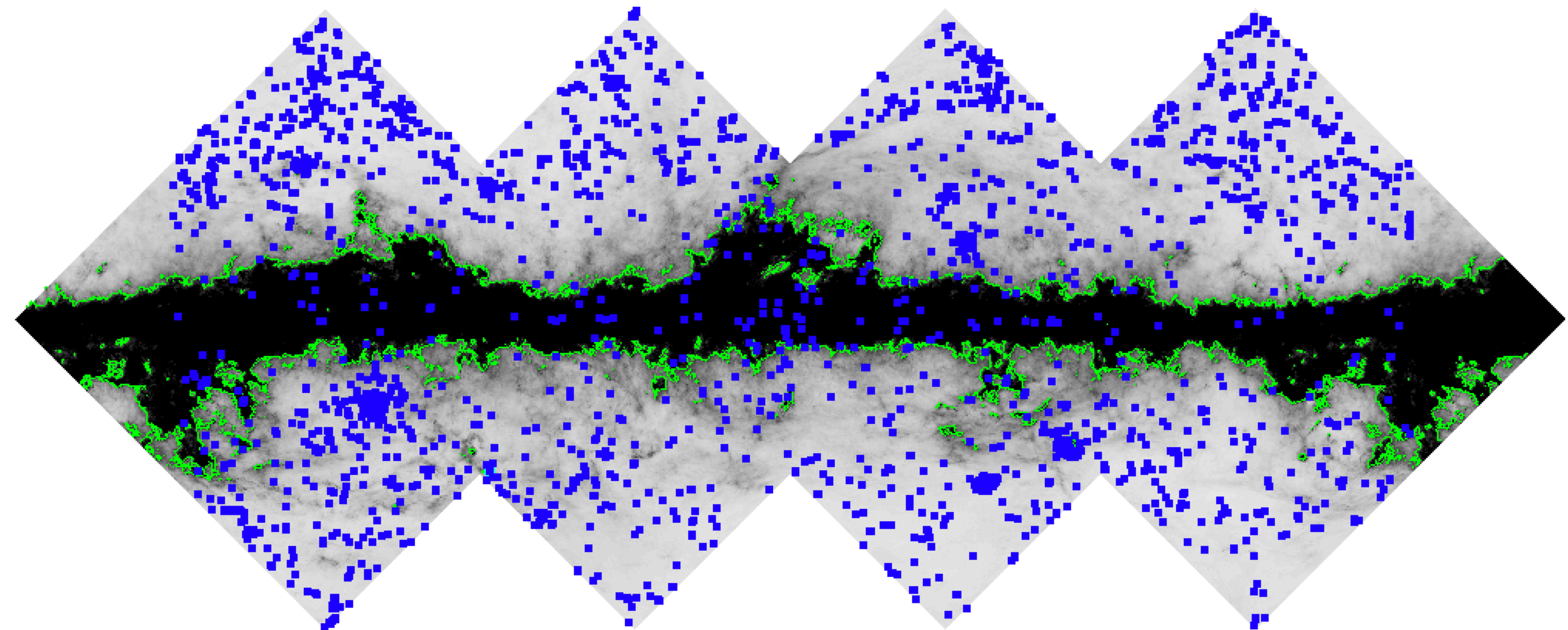
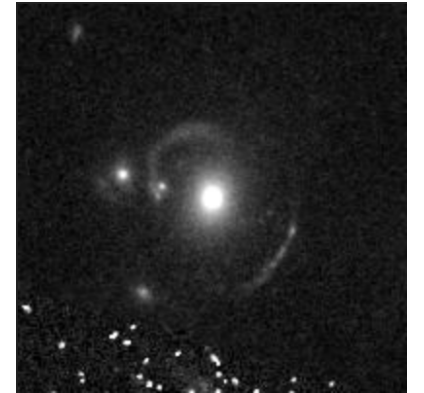
- GLAMER software
- Field and lens galaxy drawn from flagship mock
- Source from HST UDF
- Realistic PSF and noise properties
- Challenge 1.0 (Metcalf+2018)  
Challenge 2.0 (on-going, see E. Savary's talk)
- Add strong lenses in SC8 images
- Test the pipeline



# Testing the simulation with HST data

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- Hubble Source Catalog (V3) and Hubble Legacy Archive
- ~158000 objects selected ( $F814W < 22$  AB)
- 5886 ACS images in F814W ( $\sim 15$ - $18$  deg<sup>2</sup>)
- Extend the search to other ACS filters, WFPC2, WFC3-UVIS/IR



# Summary and future actions

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- Lens finding pipeline based on CNN
- Improve simulation of the training set and be prepared to make use of real Euclid data after launch
- Test the procedure on “large scale” high-resolution dataset (e.g. HST)
- Citizen science
- Spectroscopic follow-up (e.g. DESI, 4MOST, WEAVE)
- Be ready for Quick Release 1 and the first ~500 lenses