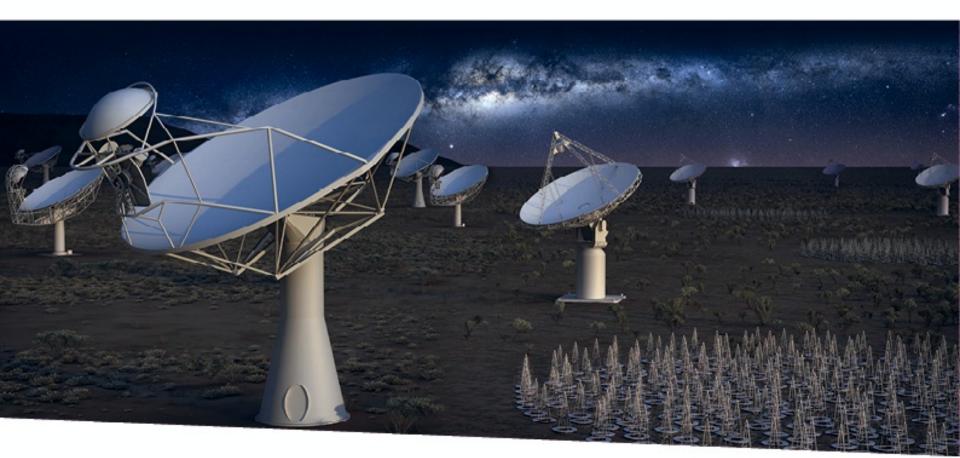
SKA Computing and Software





SQUARE KILOMETRE ARRAY

Exploring the Universe with the world's largest radio telescope

Nick Rees 18 May 2016





- Introduction
 - System overview
- Computing Elements of the SKA
 - Telescope Manager
 - Low Frequency Aperture Array
 - Central Signal Processor
 - Science Data Processor
- Conclusions

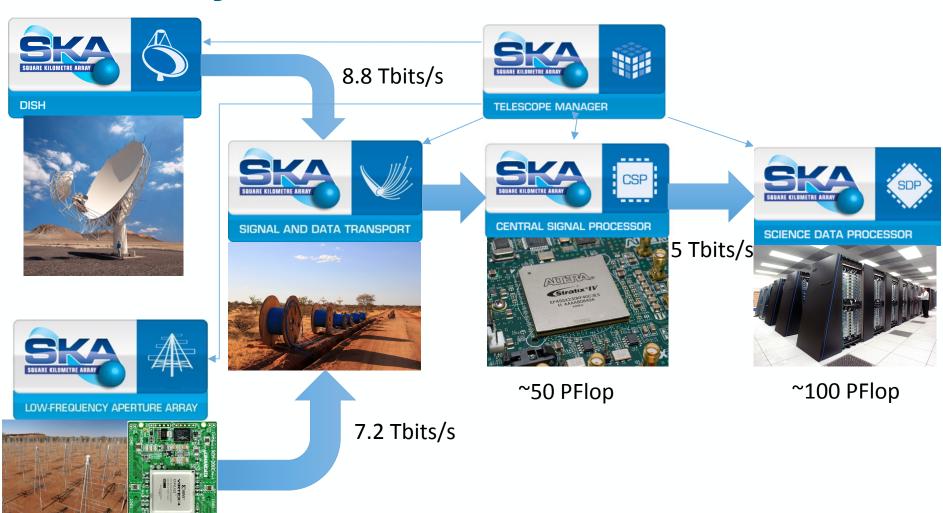
Introduction



- SKA is a software telescope
 - Very flexible and potentially easy to reconfigure
 - Major software and computing challenge
- Computing challenges are huge
 - Science Data Processor needs 100 PetaFLOPS/sec of delivered processing
 - Current Top 500 supercomputer is Tianhe-2 50 PetaFLOPS/sec
- Software challenges are also huge
 - SDP development is baselined at ~100 FTE for 6 years.
- Talk based on current status after recent PDR's
 - Designs will evolve between now and CDR.

SKA – System Overview

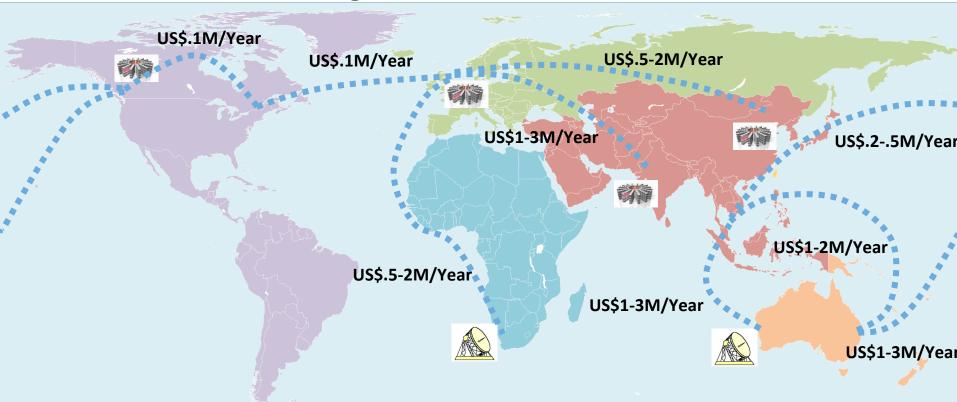






Regional Centre Network

- 10 year IRU per 100Gbit circuit 2020-2030
- Guesstimate of Regional Centres locations

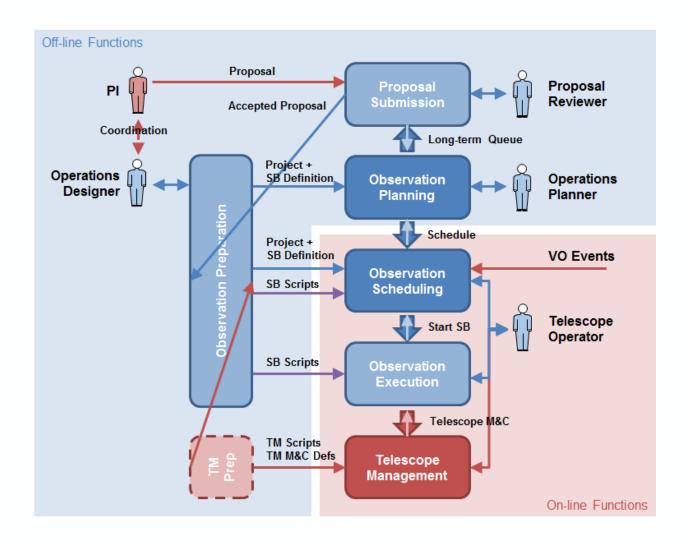




Telescope Manager

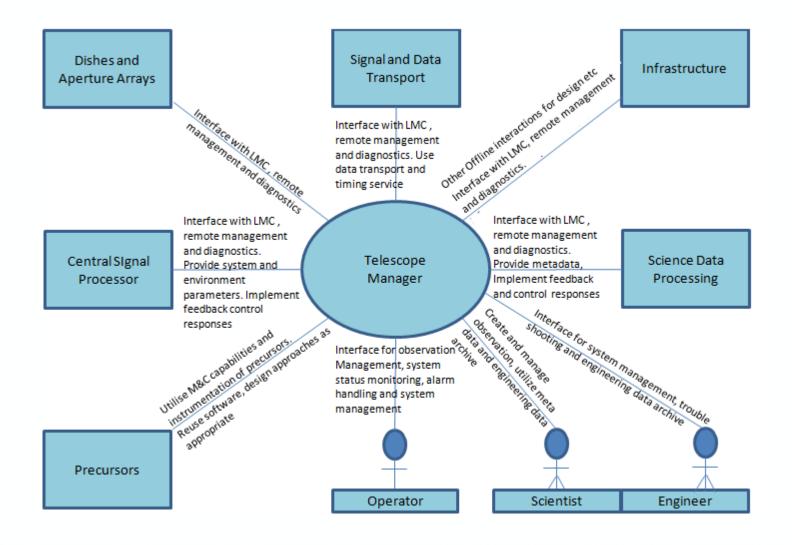


Telescope Manager Overview





Telescope Management



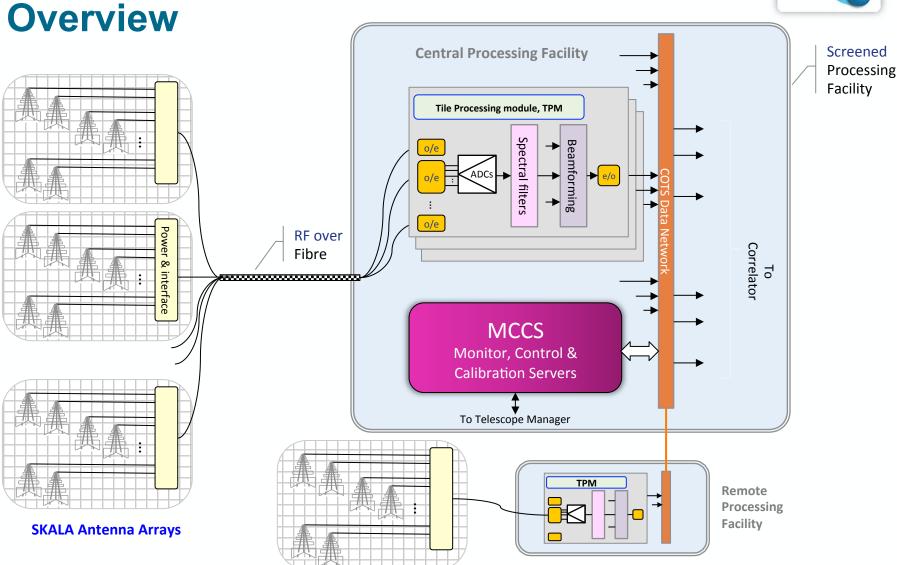




Low Frequency Aperture Array

Low Frequency Aperture Array







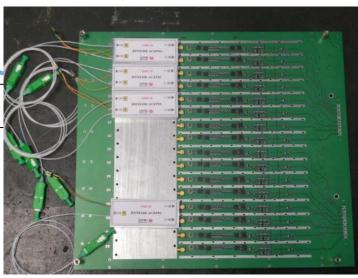
Tile Processing Modu<u>les</u>

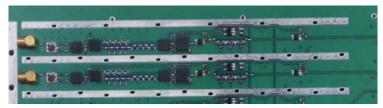


WDM optical

Analogue





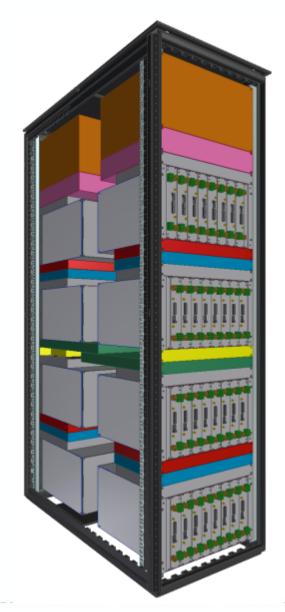


FPGA/Digital



Rack arrangement

- Double sided racks
- Water cooled using building circulation
- 64 TPMs per rack,128 in the system
- Power, Clocks and timing circulated



PSU and cooling

POWER DIST. FANS

TPM

CLOCK DIST.
PIPES DIST.
FANS

TPM

40Gb SWITCH FANS

TPM

CLOCK DIST.
PIPES DIST.
FANS

TPM



Central Signal Processor (CSP)

Central Signal Processor

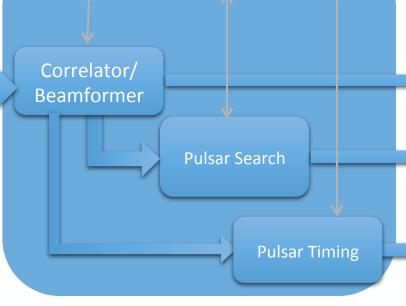
Overview















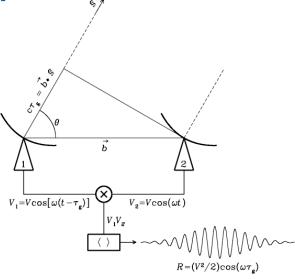
Correlator Beam Former (CBF)

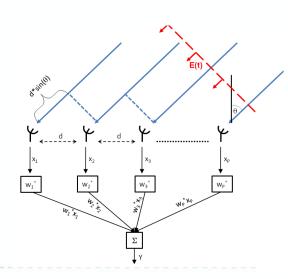
Correlator:

- Channelise signal from every dish/aperture array station in to fine frequency channels (65k)
- Cross-correlate all channels for every pair of dishes/ stations
- Cross-correlations ('visibilities') passed to SDP for imaging

Beamformer:

- Forms multiple beams within the dish/station beam
 - 1500 beams for Mid and 750 for Low
- Passes data to Pulsar Search/Timing engines/VLBI interface
- Very large amounts of real-time processing:
 - $N_{corr} \sim B(N_{dish}.log_2(N_{ch}) + N_{dish}^2) \sim PetaMAC/s$
 - $N_{BF} \sim B.N_{dish}.N_{beam} \sim few PetaMAC/s$
- Based on custom FPGA processing platforms







Pulsar Search

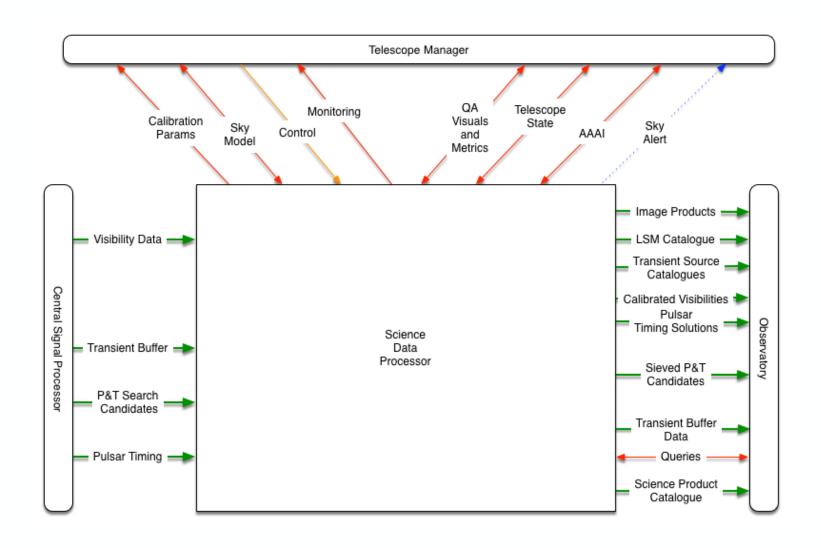
- General processing pipeline requires ~50 PFLOPS/sec.
- Baselined heterogeneous design to achieve best combination of hardware & software firmware.
- Two beams per compute node in current design.
- 250 server nodes in Australia and 750 in South Africa
- Dual redundant 10 & 1 gig networks.
- Each Node (1000 in total):
 - Low Power CPUs
 - GPUs
 - FPGA boards
 - 10 Gig inputs
 - > 1 Tbyte RAM &/or SSDs



Science Data Processor

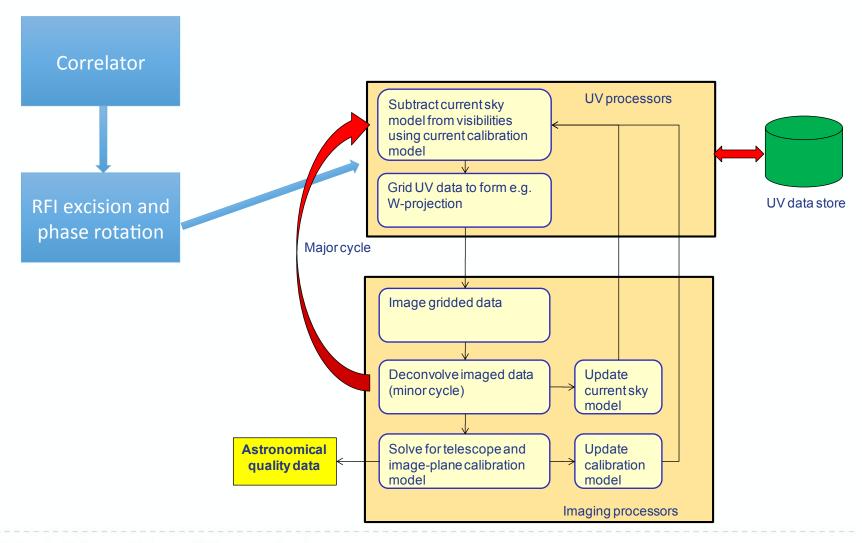


Science Data Processor Overview



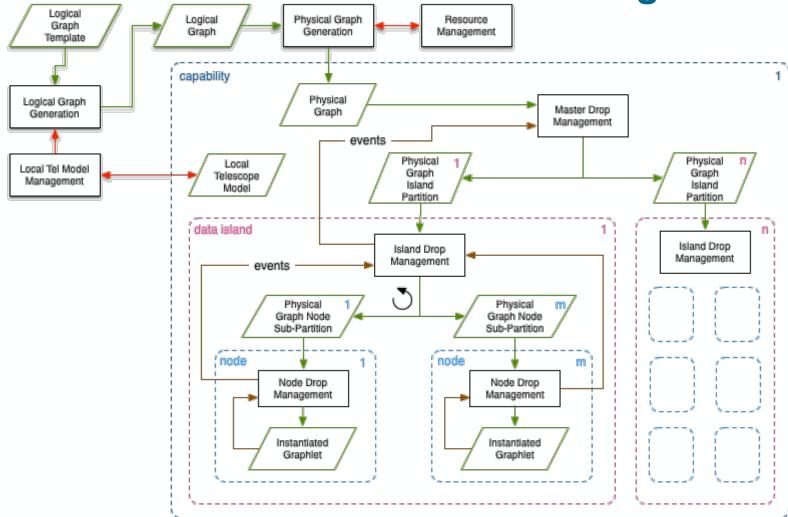


Imaging Processing Model













Computing Limitations

- Arithmetic Intensity ρ = Total FLOPS/Total DRAM Bytes
- The principal algorithms required by SDP (gridding and FFT) are typically $\rho \approx 0.5$
- Typical accelerators have an $\rho \approx 5-7$
 - For example, NVidia Pascal GPU architecture has:
 - Memory bandwidth ≈ 720 GB/sec
 - Floating point bandwidth ≈ 5,000 GFLOPS/sec
 - $\rho = 5000/720 \approx 7$
- Hence, the computational efficiency $\approx 0.5/7 \approx 10\%$
 - So, because of the bandwidth requirements, we have to buy 10 x more computing than a pure HPC system would require.
 - Unless the vendors improve the memory bandwidth...



Computing Requirements

- ~100 PetaFLOPS/sec total sustained
- ~200 PetaByte/s aggregate BW to fast working memory
- ~50 PetaByte fast working storage
- ~1 TeraByte/s sustained write to storage
- ~10 TeraByte/s sustained read from storage
 - ~ 10000 FLOPS/byte read from storage
- Current power cap proposed is ~5MW per site.



Data Management Challenges

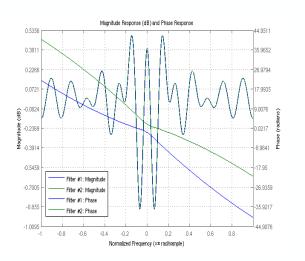
- All Top500 HPC systems have been designed for High Performance Computing (by definition).
- IDC has proposed a new term HPDA High Performance Data Analytics to reflect systems like SKA
- Must ensure the data is available when and where it is needed.
- CPU's must not be idling waiting for data to arrive
 - Data must be in fast cache when it is needed.
- Need a framework that supports this.
 - Looking at a variety of possible prototypes

Addressing Power



- Need to achieve a FLOPS/Watt an order of magnitude better than current greenest computer.
- Need a three pronged approach:

Algorithms



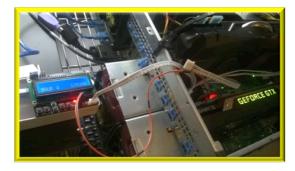
Hardware

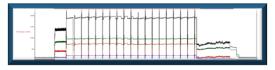


Pursue innovative approaches Look at accelerators, hosts, to cut processing times

networks and storage.

Testing





Using real algorithms and fully instrumented systems





- Budget of > €100M on manpower for software development across the whole telescope.
 - Not a task for academic programmers
 - Need professional practices for development, testing, integration and deployment.
 - Need to unify the processes across the world-wide team of developers.
 - Need world-leading expertise in a number of areas.
- Delivered system will not be static
 - SDP hardware and software will be updated periodically.
 - Key input for development will be the scientific and software community through the regional centres.



Conclusions

- SKA is a huge computational challenge
 - CSP ~ 50 Pflop, 5 MW
 - SDP ~ 100 Pflop, 5 MW
 - Tianhe-2 ~30 Pflop, 40 MW
- Traditional HPC is not a good match because the problem is bandwidth dominated.
 - SKA is seen as a key programme in global IT development
 - Showcases a major development area of High Performance Data Analysis (HPDA).
- Power is also a major driver.
- Software complexity is also beyond what has been achieved in astronomy previously.