- View through the visibilities -

Beyond CLEAN: scalable algorithms for interferometric imaging in the SKA era

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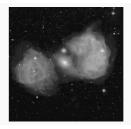
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Interferometric imaging in the SKA era

Our previous work highlighted...

- Convex optimisation and ℓ₁ minimisation Y. Wiaux, L. Jacques, G. Puy, A.M.M. Scaife, "Compressed sensing imaging techniques for radio interferometry", MNRAS, 2009
- Non-Fourier acquisition Y. Wiaux, G. Puy, Y. Boursier, P. Vandergheynst, "Spread spectrum for imaging techniques in radio interferometry", MNRAS, 2009





- Software R.E. Carrillo, J.D. McEwen, Y. Wiaux, "PURIFY: a new approach to radio-interferometric imaging", MNRAS, 2014
- Scalability A. Onose, R.E. Carrillo, A. Repetti, J.D. McEwen, J.-P. Thiran, J.-C. Pesquet, Y. Wiaux,
 "Scalable splitting algorithms for big-data interferometric imaging in the SKA era", MNRAS, submitted
- Wide-band wide-field polarised imaging, and joint calibration represent tremendous challenges...





Challenges

► Increase the resolution and sensitivity up to two orders of magnitude over current instruments

gigapixel images

huge dynamic range



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- Unprecedented amount of data to be processed
- Good reconstruction quality with scalable algorithms employing parallel and distributed processing





Inverse problem

► Measurement equation

$$y(\mathbf{u}) = \int D(\mathbf{I}, \mathbf{u}) x(\mathbf{I}) e^{-2i\pi \mathbf{u} \cdot \mathbf{I}} d^2 \mathbf{I}$$

Discretised version of the ill-posed inverse problem

$$y = \Phi x + n$$
 with $\Phi = GF$

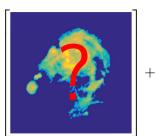
- $x \in \mathbb{R}^N_+$ the intensity image of interest
- ▶ $\Phi \in \mathbb{C}^{M \times N}$ a linear map; image domain to visibility space
- $\mathbf{y} \in \mathbb{C}^M$ the measured visibilities
- ▶ $\mathbf{G} \in \mathbb{C}^{M \times kN}$ gridding matrix modelling DDEs
- ▶ $\mathbf{F} \in \mathbb{C}^{kN \times N}$ Fourier matrix with zero padding



SKA *u*–*v* coverage



Φ







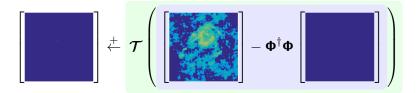
- Greedy iterative deconvolution algorithm
 - ► Select atoms associated with brightest pixel of residual image
 - ▶ Build the solution implicitly imposing sparsity in image space

$$\mathbf{x}^{(t)} = \mathbf{x}^{(t-1)} + \frac{\mathcal{T}\left(\mathbf{\Phi}^{\dagger}\left(\mathbf{y} - \mathbf{\Phi}\mathbf{x}^{(t-1)}\right)\right)$$

- Forward backward like structure
 - Forward step on the gradient direction of the ℓ_2 norm of the residual image (major cycle)
 - Backward step with non-linear sparsity enforcing operator T
 (minor cycle)



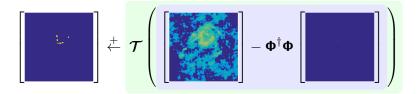




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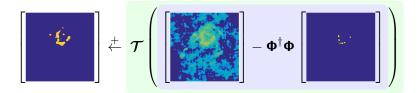




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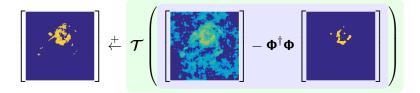




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Extreme data size

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extremely large M and N

... even more for wide-band, wide-field, polarisation data



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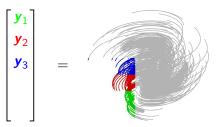
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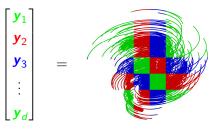
... even more for wide-band, wide-field, polarisation data

scalable algorithms, parallelisms and distributed processing

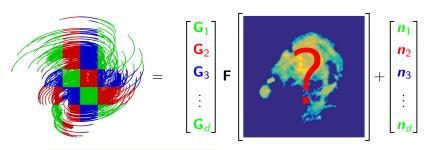






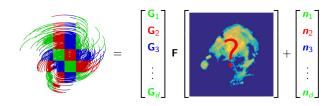






- ► Huge number of visibilities *y*
 - ▶ Distribute and process the blocks independently in parallel



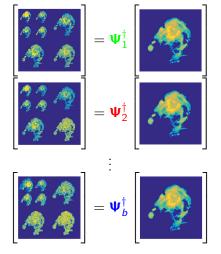


Distributed processing nodes





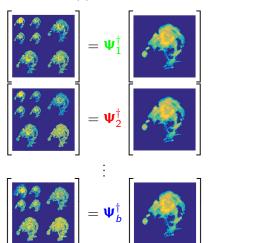




E.g. Average sparsity - a collection of wavelet bases to regularise the ill-posed problem, way beyond CLEAN.













► Split the large-scale inverse problem block wise

$$oldsymbol{y}_j = oldsymbol{\Phi}_j oldsymbol{x} + oldsymbol{n}_j$$
 with $oldsymbol{\Phi}_j = oldsymbol{\mathsf{G}}_j oldsymbol{\mathsf{F}}$

$$\mathbf{y} = \begin{bmatrix} \mathbf{y}_1 \\ \vdots \\ \mathbf{y}_d \end{bmatrix}$$
 $\mathbf{\Phi} = \begin{bmatrix} \mathbf{\Phi}_1 \\ \vdots \\ \mathbf{\Phi}_d \end{bmatrix}$

- Regularisation of the ill-posed problem
 - ▶ Sparsity constraint for *x* in a collection of wavelet bases

$$\mathbf{\Psi} = \begin{bmatrix} \mathbf{\Psi}_1 & \dots & \mathbf{\Psi}_b \end{bmatrix}$$



Convex optimisation task

$$\min_{\mathbf{x}} f(\mathbf{x}) + \gamma \sum_{i=1}^{b} l_i(\mathbf{\Psi}_i^{\dagger} \mathbf{x}) + \sum_{j=1}^{d} h_j(\mathbf{\Phi}_j \mathbf{x})$$

► Enforce positivity , sparsity and data fidelity

$$f(\mathbf{z}) = \iota_{\mathcal{C}}(\mathbf{z}), \mathcal{C} = \mathbb{R}_{+}^{N}$$

$$I_i(z) = \|z\|_1$$

$$h_j(\mathbf{z}) = \iota_{\mathcal{B}_j}(\mathbf{z}), \ \mathcal{B}_j = \{\mathbf{z} \in \mathbb{C}^{M_j} : \|\mathbf{z} - \mathbf{y}_j\|_2 \le \epsilon_j\}$$

The primal dual approach

Primal problem

$$\min_{\mathbf{x}} f(\mathbf{x}) + \gamma \sum_{i=1}^{b} l_i(\mathbf{\Psi}_i^{\dagger} \mathbf{x}) + \sum_{j=1}^{d} h_j(\mathbf{\Phi}_j \mathbf{x})$$

Dual formulation of the original convex optimisation task

$$\min_{\substack{\mathbf{u}_i \\ \mathbf{v}_j}} f^* \left(-\sum_{i=1}^b \mathbf{\Psi}_i \mathbf{u}_i - \sum_{j=1}^d \mathbf{\Phi}_j^{\dagger} \mathbf{v}_j \right) + \frac{1}{\gamma} \sum_{i=1}^b l_i^* (\mathbf{u}_i) + \sum_{j=1}^d h_j^* (\mathbf{v}_j)$$

- Primal dual algorithm
 - ▶ Alternate solving the primal problem and the dual problem
 - Converges towards a Kuhn-Tucker point





Advantages of the primal dual approach

- ► Full splitting of the operators and functions (versus ADMM)
 - ▶ No inversion of the linear operators
 - ▶ The updates are performed on the dual variables in parallel
- Interlaced and parallel CLEAN-like iteration structure
 - Forward-backward iterations are applied in parallel for all dual variables in data, sparsity, and image space
- Randomised updates on the dual variables
 - ▶ Reduces computational and memory needs per iteration
 - Requires more iterations to converge





Distributed imaging

Primal dual algorithm

$$\begin{split} & \text{given } \boldsymbol{x}^{(0)}, \tilde{\boldsymbol{x}}^{(0)}, \boldsymbol{u}_i^{(0)}, \boldsymbol{v}_j^{(0)}, \tilde{\boldsymbol{u}}_i^{(0)}, \tilde{\boldsymbol{v}}_j^{(0)}, \gamma, \tau, \sigma_i \\ & \text{repeat for } t = 1, \dots \\ & \text{generate sets } \mathcal{P} \subset \{1, \dots, b\} \text{ and } \mathcal{D} \subset \{1, \dots, d\} \\ & \boldsymbol{b}_i^{(t)} = \mathbf{M}_j \mathbf{F} \mathbf{Z} \tilde{\mathbf{x}}^{(t-1)}, \quad \forall j \in \mathcal{D} \end{split}$$

run simultaneously

 $\forall j \in \mathcal{D}$ distribute $\boldsymbol{b}_{i}^{(t)}$ and do in parallel

$$\mathbf{v}_{j}^{(t)} = \left(\mathbf{\mathcal{I}} - \mathbf{\mathcal{P}}_{\mathcal{B}_{j}}\right) \left(\mathbf{v}_{j}^{(t-1)} + \mathbf{G}_{j}\mathbf{b}_{j}^{(t)}\right) \quad \tilde{\mathbf{v}}_{j}^{(t)} = \mathbf{G}_{j}^{*}\mathbf{v}_{i}^{(t)}$$

"CLEAN v;"

end and gather $\tilde{\mathbf{v}}_{i}^{(t)}$

 $\forall i \in \mathcal{P}$ do in parallel

$$oldsymbol{u}_{i}^{(t)} = \left(\mathcal{I} - \mathcal{S}_{rac{\gamma}{\sigma_{i}}}
ight) \left(oldsymbol{u}_{i}^{(t-1)} + oldsymbol{\Psi}_{i}^{*} ilde{\mathbf{x}}^{(t)}
ight) \quad ilde{u}_{i}^{(t)} = oldsymbol{\Psi}_{i} oldsymbol{u}_{i}^{(t)}$$

$$\tilde{\boldsymbol{u}}_{i}^{(t)} = \boldsymbol{\Psi}_{i} \boldsymbol{u}_{i}^{(t)}$$

"CLEAN ũ;"

end end

$$\bar{\mathbf{x}}^{(t)} = \mathcal{P}_{\mathcal{C}}\left(\mathbf{x}^{(t-1)} - \tau\left(\sum_{i=1}^{b} \sigma_{i} \bar{\mathbf{u}}_{i}^{(t)} + \mathbf{Z}^{*} \mathbf{F}^{\dagger} \sum_{j=1}^{d} \varsigma_{j} \mathbf{M}_{j}^{*} \bar{\mathbf{v}}_{j}^{(t)}\right)\right)$$

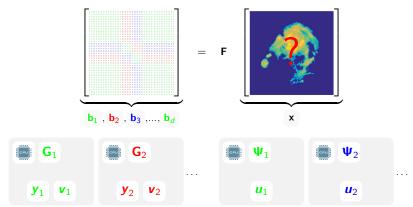
"CLEAN x"

$$\mathbf{\tilde{v}}^{(t)} - 2\mathbf{\tilde{v}}^{(t)} - \mathbf{v}^{(t-1)}$$

until convergence

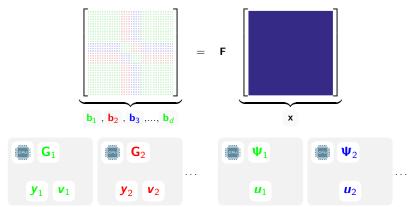




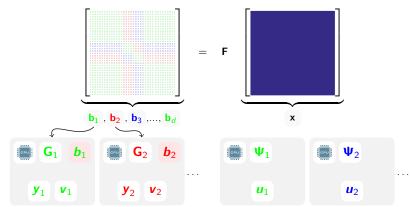






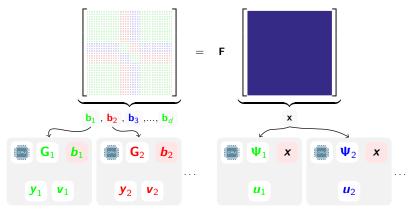






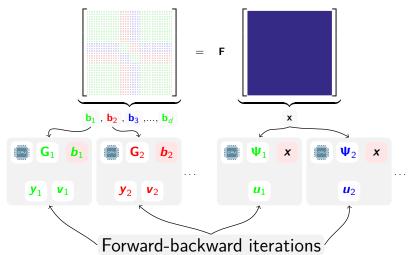
















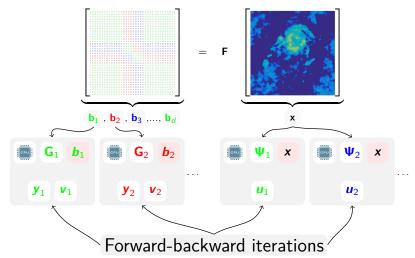
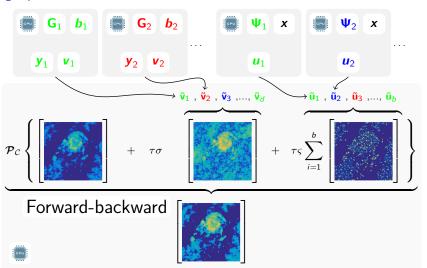






Image update







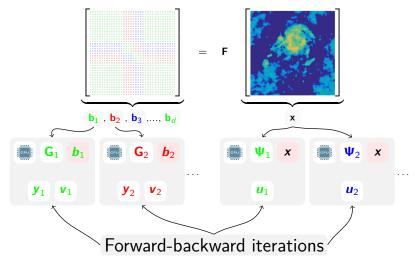
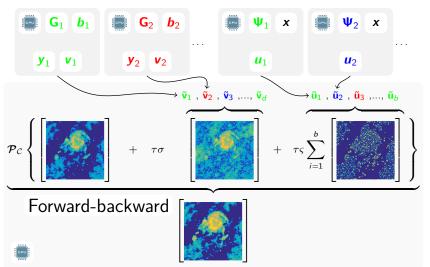






Image update







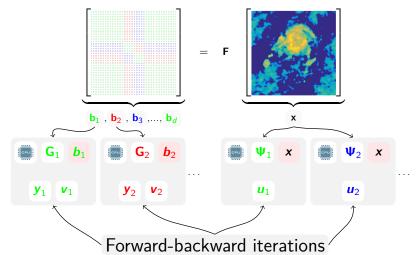
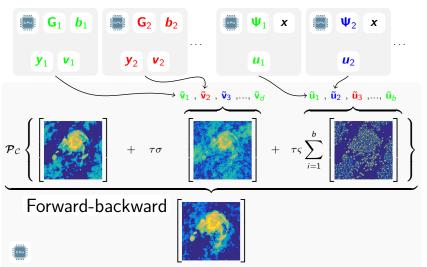






Image update







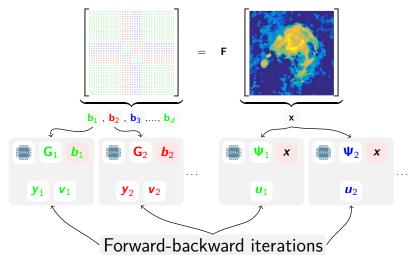
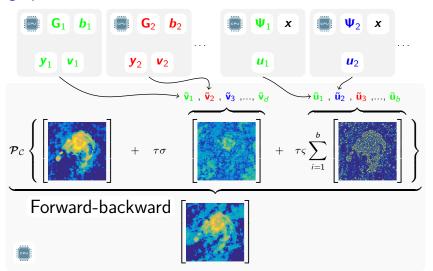






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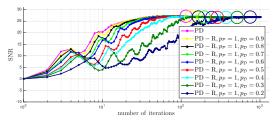
Recovery example using SKA coverage

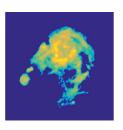


- ► SKA coverage
 - ▶ Input SNR = 40dB; Reconstruction SNR = 26dB; DR \approx 900



Randomisation





- ▶ Incomplete gaussian coverage; input SNR = 20dB
- Lower complexity per iteration but require more iterations
- Overall the computational load does not increase significantly



