

Small scale formations for the incompressible Boussinesq equation

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joint work with Alexander Kiselev and Jaemin Park

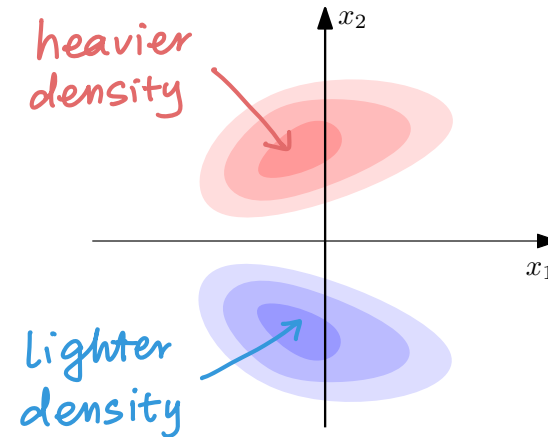
Deterministic and random features in fluids

EPFL

2D Boussinesq equation without density diffusivity

- $\rho(x, t)$: density of incompressible fluid.
- $u(x, t)$: velocity field of fluid.
- The spatial domain Ω is either the plane \mathbb{R}^2 , the torus \mathbb{T}^2 , or the strip $\mathbb{T} \times [-\pi, \pi]$.
- Throughout this talk, we consider the 2D Boussinesq equation **without density diffusivity**:

$$\begin{cases} \rho_t + u \cdot \nabla \rho = 0, \\ u_t + u \cdot \nabla u = -\nabla p - \rho e_2 + \nu \Delta u, \\ \nabla \cdot u = 0, \end{cases}$$



- We'll discuss the **viscous case** $\nu > 0$, and the **inviscid case** $\nu = 0$.
- **Goal**: In both cases, we'll prove that solution can have **small scale formation** (infinite-in-time growth of Sobolev norms) as $t \rightarrow \infty$.

The viscous case: global well-posedness and upper bounds

When $\nu > 0$, global-wellposedness of regular solutions is known:

- For $\Omega = \mathbb{R}^2$: global regularity by Hou–Li '05 in the space $(u, \rho) \in H^m \times H^{m-1}$ for $m \geq 3$, and Chae '06 in $H^m \times H^m$ for $m \geq 3$.
- For bounded Ω : global regularity by Lan–Pan–Zhao '11 in $H^3 \times H^3$, and Hu–Kukavica–Ziane '13 in $H^m \times H^{m-1}$ for $m \geq 2$.

Upper bounds for the global solution:

- Ju '17: For bounded Ω , $\|\rho\|_{H^1} \lesssim e^{Ct^2}$;
- Kukavica–Wang '20: For bounded Ω , $\|\rho\|_{H^1} \lesssim e^{Ct}$ and $\|u\|_{W^{2,p}} \leq C_p$; for \mathbb{R}^2 , $\|\rho\|_{H^1} \lesssim e^{Ct^{(1+\beta)}}$.
- Kukavica–Massatt–Ziane '21: For bounded Ω , $\|\rho\|_{H^2} \leq C_\epsilon e^{\epsilon t}$, $\|u\|_{H^3} \leq C_\epsilon e^{\epsilon t}$.

What about lower bounds?

- Note that the above estimates all deal with the **upper bounds** of solutions.
- **Question.** What about **lower bounds**? Can solutions have small scale formation as $t \rightarrow \infty$?
- Lower bound by **Brandolese–Schonbek '12**: in \mathbb{R}^2 , if ρ_0 does not have mean zero, $\|u(t)\|_{L^2} \sim (1+t)^{1/4}$. (This is due to potential energy converting to kinetic energy, and does not imply growth in higher derivatives)
- We are not aware of any examples of infinite-in-time growth of $\|\rho(t)\|_{H^m}$ in the literature.

Small scale formation in the viscous case

Theorem (Kiselev–Park–Y. '22, preprint)

Let $\nu > 0$, $\Omega = \mathbb{T}^2$. If the smooth initial data (ρ_0, u_0) satisfies the following

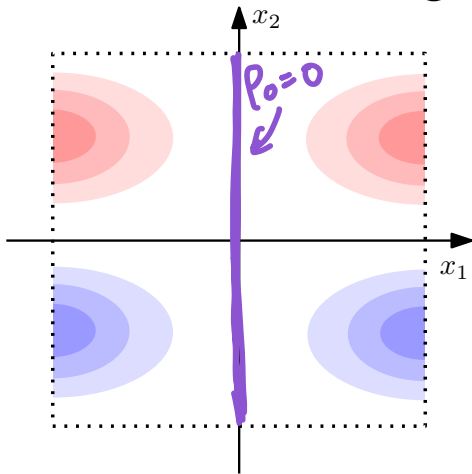
- **Symmetry assumptions:** ρ_0 is even-odd, u_{01} is odd-even, u_{02} is even-odd.
- **Sign assumptions:** $\rho_0 \geq 0$ for $x_2 \geq 0$, and $\rho_0 = 0$ on the x_2 -axis.

Then the global-in-time smooth solution satisfies

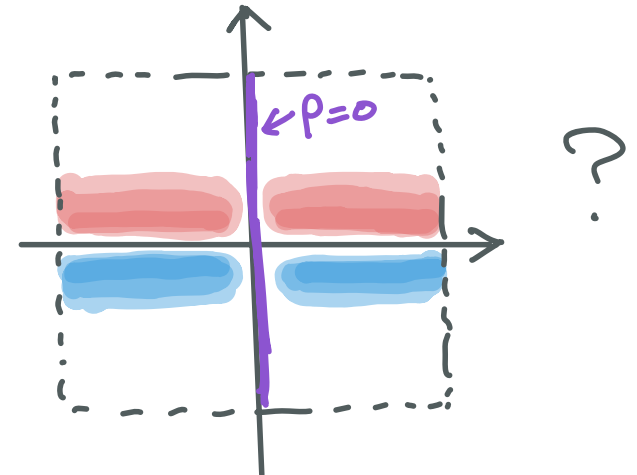
$$\limsup_{t \rightarrow \infty} t^{-\frac{1}{6}} \|\rho(t)\|_{\dot{H}^1} = +\infty.$$

Preserved for all time!

Remark: Under these assumptions one can show $\|\rho(t)\|_{\dot{H}^1}$ has a refined **sub-exponential upper bound** $\exp(Ct^\alpha)$ for some $\alpha \in (0, 1)$, so the growth is somewhere between algebraic and sub-exponential.



Asymptotics as $t \rightarrow \infty$?
(Still open!)



Evolution of potential energy

- Define the **potential energy** $E_P(t) := \int_{\mathbb{T}^2} \rho x_2 dx$, and **kinetic energy** $E_K(t) := \int_{\mathbb{T}^2} |u|^2 dx$.

- It's well-known that the **total energy** is **decreasing in time**:

$$\frac{d}{dt}(E_P(t) + E_K(t)) = -\nu \|\nabla u(t)\|_{L^2}^2.$$

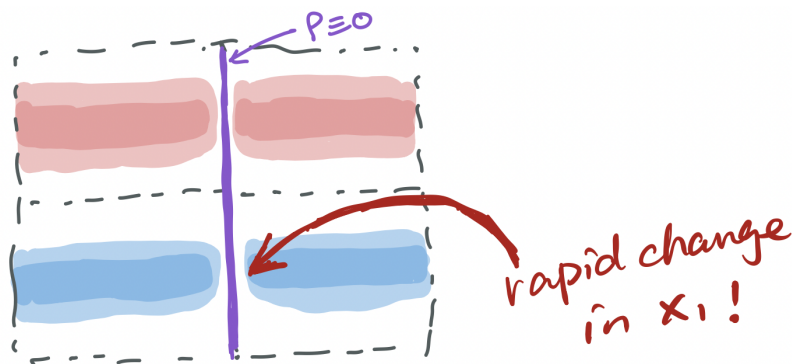
This implies that $\int_0^\infty \|\nabla u(t)\|_{L^2}^2 dt < C(\nu, \rho_0, u_0)$.

- Since the two equations are coupled by the gravity force, we'll track the evolution of potential energy $E_P(t)$ itself.
- A quick computation gives $\frac{d}{dt} E_P(t) = \int_{\mathbb{T}^2} \rho u_2 dx$, which is uniformly bounded.

- Let's take one more time derivative:

$$\frac{d^2}{dt^2} E_P(t) = \underbrace{\sum_{i,j=1}^2 \int_{\mathbb{T}^2} \overbrace{((-\Delta)^{-1} \partial_2 \rho)}^{\text{bounded}} \partial_i u_j \partial_j u_i dx}_{=:A(t)} - \underbrace{\nu \int_{\mathbb{T}^2} \nabla \rho \cdot \nabla u_2 dx}_{=:B(t)} - \|\partial_1 \rho\|_{\dot{H}^{-1}}^2.$$

- Since $\int_0^\infty \|\nabla u(t)\|_{L^2}^2 dt < \infty$, this implies $\int_0^\infty A(t) dt < \infty$.
- Suppose $\limsup_{t \rightarrow \infty} \|\nabla \rho\|_{L^2} < \infty$, we have $\int_0^\infty B(t) dt \lesssim t^{1/2}$.
- This implies $\int_0^\infty \|\partial_1 \rho\|_{\dot{H}^{-1}}^2 dt \lesssim t^{1/2}$, so $\|\partial_1 \rho\|_{\dot{H}^{-1}}^2$ needs to decay to zero like $t^{-1/2}$ as $t \rightarrow \infty$.
- Key observation (by a Fourier argument):** If $\|\partial_1 \rho(t)\|_{\dot{H}^{-1}} \ll 1$ and $\rho \equiv 0$ on x_2 axis, we have $\|\rho\|_{\dot{H}^1} \gg 1$. More precisely, $\|\rho\|_{\dot{H}^1} \gtrsim \|\partial_1 \rho(t)\|_{\dot{H}^{-1}}^{-1}$.



- This contradicts our assumption $\limsup_{t \rightarrow \infty} \|\nabla \rho\|_{L^2} < \infty$. (A more careful argument gives us algebraic growth in time).

Inviscid 2D Boussinesq equation

- In the inviscid case $\mu = 0$, let us work with the variables ρ and vorticity ω :

$$\begin{cases} \rho_t + u \cdot \nabla \rho = 0, \\ \omega_t + u \cdot \nabla \omega = -\partial_1 \rho, \end{cases}$$

where u can be recovered from the Biot-Savart law $u = \nabla^\perp (-\Delta)^{-1} \omega$.

- Whether smooth initial data can lead to a blow-up in \mathbb{T}^2 or \mathbb{R}^2 is an outstanding open question.
- It is well-known that away from the axis of symmetry, the 3D axisymmetric Euler equation is closely related to 2D Boussinesq:

$$\begin{cases} D_t(ru^\theta) = 0, \\ D_t\left(\frac{\omega^\theta}{r}\right) = r^{-4} \partial_z (ru^\theta)^2, \end{cases}$$

where $D_t := \partial_t + u^r \partial_r + u^z \partial_z$ is the material derivative, and (u^r, u^z) can be recovered from ω^θ / r by a similar Biot-Savart law.

Blow-up for inviscid 2D Boussinesq and 3D Euler

In the **presence of boundary**, or for **non-smooth initial data**, there are many exciting developments on finite-time blow-up:

- **Luo–Hou '14**: convincing numerical evidence for blow-up at the boundary for 3D axisymmetric Euler
- **Elgindi–Jeong '20**: blow-up in domain with a corner
- **Elgindi '21**: blow-up for $C^{1,\alpha}$ solutions for 3D Euler
- **Chen–Hou '21**: blow-up for $C^{1,\alpha}$ solutions with boundary
- **Wang–Lai–Gómez-Serrano–Buckmaster '22**: numerics for approximate self-similar blow-up solution using physics-informed neural networks.
- **Chen–Hou '22**: stable nearly self-similar blowup for smooth solutions (combination of analysis + computer-assisted estimates)

Question: Can one construct solutions with infinite-in-time growth for more general class of initial data?

Infinite-in-time growth in a strip

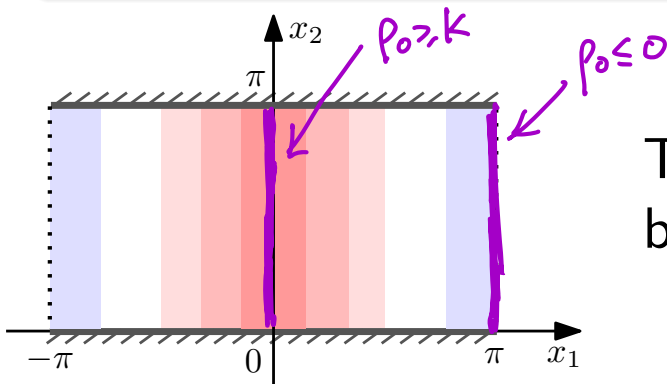
Theorem (Kiselev–Park–Y. '22, preprint)

Let $\Omega = \mathbb{T} \times [0, \pi]$. Let $\rho_0 \in C^\infty(\Omega)$ be even in x_1 , and $\omega_0 \in C^\infty(\Omega)$ be odd in x_1 , with $\int_{[0, \pi] \times [0, \pi]} \omega_0 dx \geq 0$. Assume that there exists $k_0 > 0$ such that $\rho_0 \geq k_0 > 0$ on $\{0\} \times [0, \pi]$, and $\rho_0 \leq 0$ on $\{\pi\} \times [0, \pi]$. Then the solution satisfies the following during its lifespan:

$$\|\omega(t)\|_{L^p(\Omega)} \gtrsim t^{3 - \frac{2}{p}},$$

$$\|u(t)\|_{L^\infty(\Omega)} \gtrsim t,$$

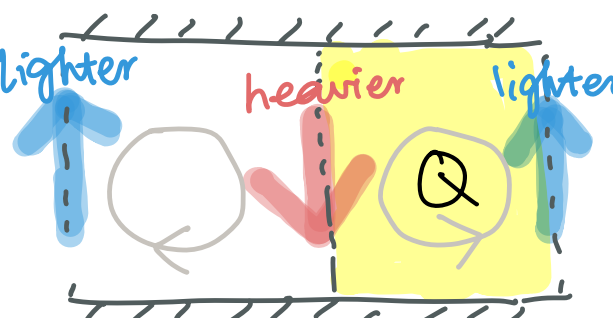
$$\sup_{\tau \in [0, t]} \|\nabla \rho(\tau)\|_{L^\infty(\Omega)} \gtrsim t^2.$$



The proof is a soft argument, based on an interplay between various monotone and conservative quantities.

Monotonicity of vorticity integral

- Let Q be the right half of the strip. Simple but useful observation:



$$\begin{aligned}
 \frac{d}{dt} \int_Q \omega dx &= \int_Q -u \cdot \nabla \omega dx - \int_Q \partial_1 \rho dx \\
 &= \int_0^\pi \underbrace{\rho(0, x_2, t)}_{\geq k_0} dx_2 - \int_0^\pi \underbrace{\rho(\pi, x_2, t)}_{\leq 0} dx_2 \\
 &\geq k_0 \pi.
 \end{aligned}$$

- Since $\int_{\partial Q} u \cdot dl = \int_Q \omega dx \geq k_0 \pi t$, we have $\|u(t)\|_{L^\infty}$ grows at least linearly.
- On the other hand, $\|u\|_{L^2}$ is bounded for all times by energy conservation.
- Combining the **boundedness of $\|u\|_{L^2(Q)}$** and **linear growth of $\int_{\partial Q} u \cdot dl$** , we know u must change rapidly in a small neighborhood of ∂Q , leading to super-linear growth of ∇u (and ω).

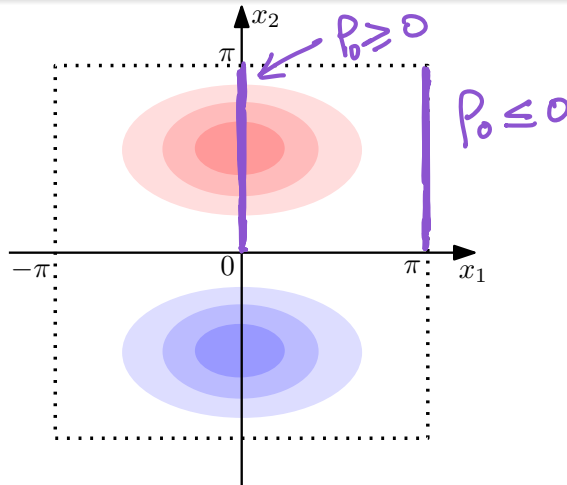
Infinite-in-time growth in \mathbb{T}^2

- To our best knowledge, there has been no blow-up / infinite-in-time growth results in \mathbb{T}^2 .
- In \mathbb{T}^2 , we obtain infinite-in-time growth for a large class of initial data satisfying certain symmetry/sign conditions:

Theorem (Kiselev–Park–Y. '22, preprint)

Let $\rho_0 \in C^\infty(\mathbb{T}^2)$ be even-odd, and $\omega_0 \in C^\infty(\mathbb{T}^2)$ be odd-odd. Assume $\rho_0 \geq 0$ on $\{0\} \times [0, \pi]$ with $k_0 := \sup_{x_2 \in [0, \pi]} \rho_0(0, x_2) > 0$, and $\rho_0 \leq 0$ on $\{\pi\} \times [0, \pi]$. Then the solution satisfies the following during its lifespan:

$$\sup_{\tau \in [0, t]} \|\nabla \rho(\tau)\|_{L^\infty(\mathbb{T}^2)} \gtrsim t^{1/2}. \quad (1)$$



3D axisymmetric Euler in an annular cylinder

Using a similar idea, we obtain infinite-in-time growth for the 3D axisymmetric Euler equation in an annular cylinder

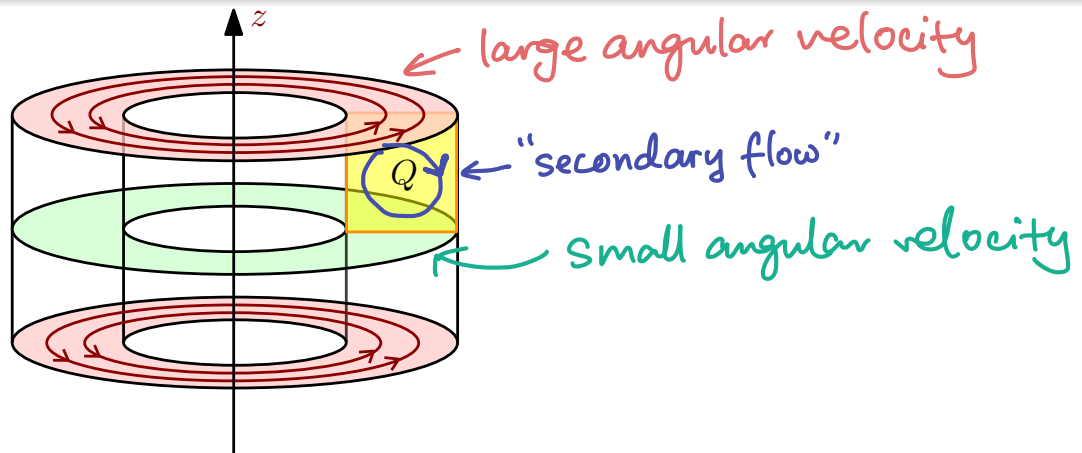
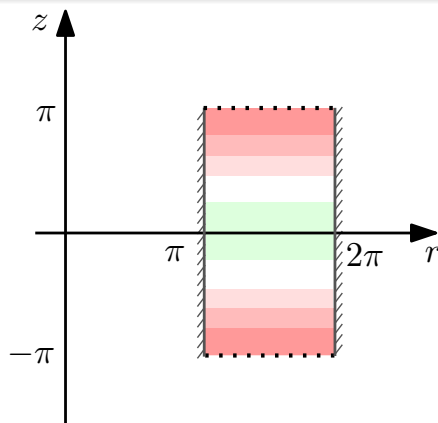
$$\Omega = \{(r, \theta, z) : r \in [\pi, 2\pi]; \theta \in \mathbb{T}, z \in \mathbb{T}\}.$$

Theorem (Kiselev–Park–Y. '22, preprint)

Let $u_0^\theta \in C^\infty(\Omega)$ be even in z , $\omega_0^\theta \in C^\infty(\Omega)$ odd in z , with $\int_0^\pi \int_\pi^{2\pi} \omega_0^\theta dr dz \geq 0$. Assume there exists $k_0 > 0$ such that $u_0^\theta \geq k_0 > 0$ on $z = \pi$, and $|u_0^\theta| \leq \frac{1}{8}k_0$ on $z = 0$. Then the solution to axisymmetric 3D Euler satisfies

$$\|\omega^\theta(t)\|_{L^p(\Omega)} \gtrsim t^{3-\frac{2}{p}} \quad \text{and} \quad \|u(t)\|_{L^\infty(\Omega)} \gtrsim t$$

during the lifespan of the solution.



Thank you for your attention!



Boussinesq eq
at Singapore
airport!